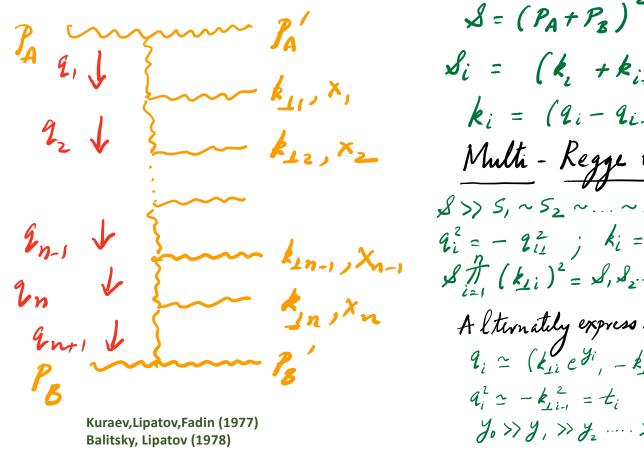
Classicalization and unitarization of wee partons in QCD and gravity: The CGC-Black Hole correspondence

Raju Venugopalan Brookhaven National Laboratory

Talk Outline

- \square 2 \rightarrow N amplitudes in the Regge limit of QCD: the BFKL equation
- Classicalization and perturbative unitarization in QCD: the Color Glass Condensate
- \square 2 \rightarrow N amplitudes in the Regge limit of Gravity: the Lipatov-ACV approach and the double copy
- Classicalization and unitarization in Gravity: The Black Hole N Portrait
- ☐ The Bekenstein-Hawking bound and unitarization in the BHNP
- ☐ The CGC/BHNP correspondence
- Coda: Gravitation memory and color memory

How does one compute multiparticle production in pQCD at high energies?



$$S = (P_{A} + P_{B}) = 2P_{A} \cdot P_{B}$$

$$S_{i} = (k_{i} + k_{i+1})^{2} = 2k_{i} \cdot k_{i+1}$$

$$k_{i} = (q_{i} - q_{i+1})_{j} \cdot k_{0} = P_{A}'_{i} \cdot k_{n+1} = P_{B}'_{i}$$

$$Multi - Regge limit:$$

$$S >> S, \sim S_{2} \sim ... \sim S_{n+1} >> q_{1}^{2} q_{2}^{2} \sim ... q_{n+1},$$

$$q_{i}^{2} = -q_{i,1}^{2}_{i} \cdot k_{i} = q_{i}P_{B} + P_{i}P_{A} + k_{1}i$$

$$S_{i=1}^{n} (k_{1}i)^{2} = S_{i}S_{2}...S_{n+1}^{n} \cdot k_{1}i^{2}P_{A} = k_{1}i \cdot P_{B} = 0$$

$$A \text{ lturnately express in terms of rapidities}$$

$$q_{i} \simeq (k_{1}i e^{3i}, -k_{1}i_{1}e^{-3i_{1}}, -k_{1}i_{1})$$

$$q_{i}^{2} \simeq -k_{1}i_{1}^{2} = t_{i}$$

$$y_{0} >> y_{1} >> y_{2} \cdot ... >> y_{n+1}^{n}; y_{i} \sim ln(x_{i})$$

To build in real and virtual corrections to all orders in α_S , first focus on one rung of 2 \rightarrow N+2 ladder

Building blocks: Real Emission Lipator vertex $((q_{i+1}, q_i) = -q_{ji+1} - q_{ji}$ Non-brul - gange invariant! Ward identity: ki C = 0 $(S,t) \sim 2S \frac{1}{t} \subset (2z,2,1) \frac{1}{t_2}$ BFKL pedagogical review: DelDuca, hep-ph/9503226

Building Blocks: virtual corrections



Reggeization angatz:

In general

$$\alpha(t) = \tilde{g}_{S}^{2} \alpha^{(1)}(t) + \tilde{g}_{S}^{4} \alpha^{(2)}(t) + \mathcal{O}(\tilde{g}_{S}^{6})$$

V. Fadin, hep-ph/9807528

with
$$\alpha(t) \propto \ln \frac{t}{t}$$
 with $\alpha(t) \propto \ln \frac{-t}{u^2}$

Infrared cut off

$$\equiv \frac{1}{t_i} \Rightarrow \frac{1}{t_i} e^{\alpha(t_i)(y_i - y_i)}$$

Double log structure:

Sudakor form factor

> infraved sensitive

> 2 > 2 amp vanishes for m >

Building Blocks: virtual corrections

$$k = \frac{1}{t} \Rightarrow \frac{1}{t} \ln \frac{s}{s} \propto (t)$$

$$with \propto (t) \propto \ln \frac{t}{m^2}$$

$$Infrared \iff m^2$$

Reggeization angatz:

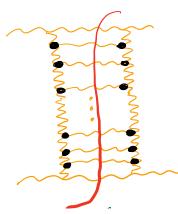
Parenthetical remark:

To this 2-loop order, this Regge trajectory can be obtained from the "cusp" anomalous dimension of the product of two Wilson lines: "Infrared factorization"

Korchemsky, Korchemskaya, hep-ph/9607229

Putting together real and virtual contributions in this leading logs in x approximation: $(\alpha_S Y)^n$

BFKL Pomeron



Im
$$A(x,t) \propto \sum_{n=0}^{\infty} (x_{3}C_{T})^{n+2}$$
 $X \int \frac{n}{11} dy_{i} \int d^{2}q_{j1}$
 $L=1 \int \frac{1}{4\pi} \int d^{2}l \int (2\pi)^{2}$
 $X = 2i \int \frac{1}{2\pi} \int e^{(y_{2}-y_{2})(x/l_{2}) + x/l_{2}} \int e^{(y_{2}-y_{2})(x/l_{2}) + x/l_{2}}$
 $X = 1 \int e^{(y_{2}-y_{2})(x/l_{2}) + x/l_{2}} \int e^{(y_{2}-y_{2})(x/l_{2})} \int e^{(y_{2}-y_{2}$

C_T is color factor

Phase space factors

Reggeized propagators on both sides of cut

Product of Lipatov vertices

$$F_{e}(t) = -2it(4\pi d_{s})^{2}G_{+}^{2} \times Phase space}$$

$$\times \frac{1}{(-2\alpha_{s}G_{+})} C_{n}C^{n}$$

$$+_{i}t'_{i} \frac{1}{1-1-\alpha'(t_{i})-\alpha'(t_{2})}$$

$$\times \frac{1}{(2\pi)^{2}} \frac{1}{q_{i_{1}}^{2}(q_{1}q_{i_{1}}^{2})}$$

$$= \frac{1}{(2\pi)^{2}} \frac{1}{q_{i_{1}}^{2}(q_{1}q_{i_{1}}^{2})}$$

$$= \frac{1}{(2\pi)^{2}} \frac{1}{q_{i_{1}}^{2}(q_{1}q_{i_{1}}^{2})}$$

Defining
$$F_{\varrho}(t) \propto \int \frac{d^2_{H_1}}{(2\pi)^2} \frac{\int_{\varrho} (\ell_{L_1} t)}{q^2_{L_1} (\ell_{L_1})^2_{L_1}}$$

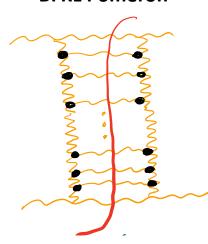
Obtain BFKL integral eqn.

$$[l-1-x(t)-x(t')]f_{\ell}(t,t_{1})$$

$$= 1-2x_{5}G\int_{(2\pi)^{2}}^{d^{2}f_{2}}\frac{C_{m}C^{m}(4,\xi_{1})}{q_{2}^{2}(2-\xi_{2})_{2}^{2}f_{\ell}^{2}t'}$$

Putting together real and virtual contributions in this leading logs in x approximation: $(\alpha_S Y)^n$

BFKL Pomeron



Im
$$f(x,t) \propto \sum_{i=1}^{\infty} (x_{i}x_{i}x_{j})^{2}$$
 $X \int_{i=1}^{n} \frac{dy_{i}}{4\pi} \int_{i=1}^{n} \frac{d^{2}q_{j1}}{(2\pi)^{2}}$
 $X = \sum_{i=1}^{n+1} \frac{dy_{i}}{4\pi} \int_{i=1}^{n+1} \frac{d^{2}q_{j1}}{(2\pi)^{2}}$

Product of Lipatov vertices

Inverse Laplace Transform:

Im A (8,t) = g dl e g f (t)

$$= 2 \operatorname{Im} A(S, t=0)$$

$$= S \operatorname{with} \lambda = 4 d_{S} N c \ln_{e} 2$$

$$\simeq 0.5 \text{ for } d_{5} = 0.2$$

$$\simeq 0.5$$
 for $d_5 = 0.2$

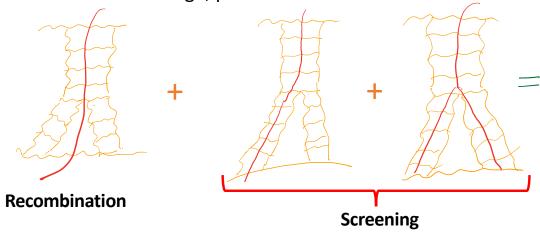
Note:

Real and virtual corrections combine to cancel **Infrared divergence!**

Strong violation of Froissart bound

NLO BFKL : $\lambda \approx 0.3$

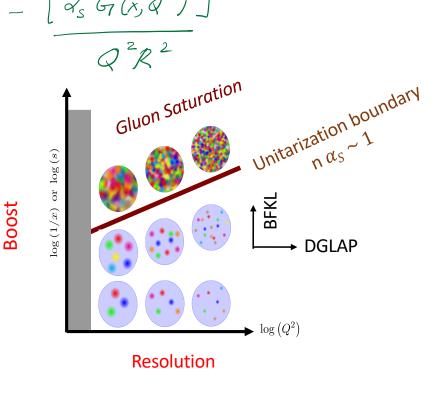
When N becomes large, partons start to recombine and screen – "shadowing"



Rapid growth in gluon occupancy with BFKL countered by "higher twist" recombination and screening, creating state of maximal occupancy

$$N = \frac{1}{\alpha_S}$$
 on saturation line Q = Q_S(x)

Classicalization! Yang-Mills non-linear dynamics of glue



Semi-classical picture: Given a non-perturbative distribution of classical color sources coupled to overoccupied background glue, compute small quantum fluctuations in a small rapidity window, and absorb these into the classical source distribution for the next window in rapidity – *interate* (RG)

$$\langle \mathrm{d}\sigma_{\mathrm{LO}} \rangle = \int [\mathcal{D}\rho_{A}] \, W_{\Lambda_{0}^{-}}[\rho_{A}] \, \mathrm{d}\hat{\sigma}_{\mathrm{LO}}[\rho_{A}] \, .$$

$$\langle \mathrm{d}\sigma_{\mathrm{LO}} + \delta\sigma_{\mathrm{NLO}:1} \rangle = \int [\mathcal{D}\rho_{A}] \Big\{ \Big(1 + \ln(\Lambda_{1}^{-}/\Lambda_{0}^{-})\mathcal{H}_{\mathrm{LO}} \Big) W_{\Lambda_{0}^{-}}[\rho_{A}] \Big\} \, \mathrm{d}\hat{\sigma}_{\mathrm{LO}}[\rho_{A}]$$

$$\Big(1 + \ln(\Lambda_{1}^{-}/\Lambda_{0}^{-})\mathcal{H}_{\mathrm{LO}} \Big) W_{\Lambda_{0}^{-}}[\rho_{A}] = W_{\Lambda_{1}^{-}}[\rho_{A}]$$

Semi-classical picture: Given a non-perturbative distribution of classical color sources coupled to overoccupied background glue, compute small quantum fluctuations in a small rapidity window, and absorb these into the classical source distribution for the next window in rapidity – *interate* (RG)

$$\langle d\sigma_{LO} \rangle = \int [\mathcal{D}\rho_{A}] W_{\Lambda_{0}^{-}}[\rho_{A}] d\hat{\sigma}_{LO}[\rho_{A}]. \qquad \langle d\sigma_{LO} + \delta\sigma_{NLO:1} \rangle = \int [\mathcal{D}\rho_{A}] \left\{ \left(1 + \ln(\Lambda_{1}^{-}/\Lambda_{0}^{-})\mathcal{H}_{LO}\right) W_{\Lambda_{0}^{-}}[\rho_{A}] \right\} d\hat{\sigma}_{LO}[\rho_{A}]$$

$$= \begin{pmatrix} (1 + \ln(\Lambda_{1}^{-}/\Lambda_{0}^{-})\mathcal{H}_{LO}) W_{\Lambda_{0}^{-}}[\rho_{A}] + W_{\Lambda_{1}^{-}}[\rho_{A}] \end{pmatrix}$$

$$= \begin{pmatrix} (1 + \ln(\Lambda_{1}^{-}/\Lambda_{0}^{-})\mathcal{H}_{LO}) W_{\Lambda_{0}^{-}}[\rho_{A}] + W_{\Lambda_{1}^{-}}[\rho_{A}] \end{pmatrix}$$

$$= \begin{pmatrix} (P) \int_{Q} (P) \int_{Q}$$

$$U = P \exp \left(i \int_{y}^{\infty} dy' \frac{\rho(x_{t}, y')}{\nabla_{t}^{2}} \right)$$
 path ordered exponential in adjoint rep. sources (likewise V in fundamental rep.) capturing "coherent" multiple scattering to all orders in ρ/∇_{t}^{2}

Semi-classical picture: Given a non-perturbative distribution of color sources coupled to classical (overoccupied) background, compute small quantum fluctuations in a small rapidity window, and absorb these into the source distribution for the next window in rapidity – *interate* (RG)

$$\langle d\sigma_{LO} \rangle = \int [\mathcal{D}\rho_{A}] W_{\Lambda_{0}^{-}}[\rho_{A}] d\hat{\sigma}_{LO}[\rho_{A}] . \qquad \langle d\sigma_{LO} + \delta\sigma_{NLO:1} \rangle = \int [\mathcal{D}\rho_{A}] \left\{ \left(1 + \ln(\Lambda_{1}^{-}/\Lambda_{0}^{-})\mathcal{H}_{LO} \right) W_{\Lambda_{0}^{-}}[\rho_{A}] \right\} d\hat{\sigma}_{LO}[\rho_{A}]$$

$$= - + \frac{\left\{ \left\{ \frac{1}{2} + \frac{1}{2} \right\} \right\} \left\{ \left(\frac{1}{2} + \frac{1}{2}$$

Parenthetical remark: In Lipatov's Reggeon EFT, these are quark-quark-reggeon and gluon-gluon-reggeon vertices respectively

Bondarenko, Lipatov, Pozdynyakov, Prygarin, arXiv:1708.05183
Hentschinski, arXiv:1802.06755

Consider quark scattering on dense target (forward physics)

Amp. × c.c. $\rightarrow 1 - \frac{1}{N_c} \langle \mathcal{F}_c(V(K_L)V(Y_L)) \rangle$ BUILDING BLOCK OF HIGH ENERGY &CD

if x,>> x2>> x3

Balitsky-Kovchegov equation: RG evolution of dipole correlator

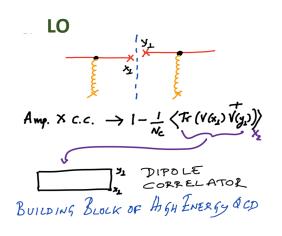
 $\frac{ds}{dy} = K_{BFKL} \otimes (5-55)$

These soft tremsstrahlung contributions above ted in RG evolution of DIPOLE CORRECATOR

 $\leq_{y_{i}} = \frac{1}{N_{c}} \langle \mathcal{T}_{r} \left(V_{x_{1}} V_{x_{2}}^{\lambda} \right) \rangle_{y_{i}}$ $\longrightarrow \frac{1}{N_{c}} \langle \mathcal{T}_{r} \left(V_{x_{1}} V_{x_{2}}^{\lambda} \right) \rangle_{y_{2}}$

SOFT BREMSSTRAHLUNG + MULTIPLE SCATTERING = SHADOWING

Consider quark scattering on dense target (forward physics)



NLO $i \neq X_1 >> X_2 >> X_3$

These soft bremsstrahlung, contributions absorbed in RG evolution of DIPOLE CORRECATOR

BK wintainings cross-section

when
$$S \to 0$$
: "Black disc"

This occurs when $N \sim 1/ds$

when dipole size $T_1 = X_1 \cdot Y_2 = \frac{1}{Q_s}$

in S_1/Z_2 , obtain BFKL eqn!

$$\leq_{y_{i}} = \frac{1}{N_{c}} \left\langle \mathcal{T}_{r} \left(V_{x_{1}} V_{x_{2}}^{\dagger} \right) \right\rangle_{y_{i}}$$

$$\longrightarrow \frac{1}{N_{c}} \left\langle \mathcal{T}_{r} \left(V_{x_{1}} V_{x_{2}}^{\dagger} \right) \right\rangle_{y_{2}}$$

Consider double in chisive now $PA \rightarrow 2g + X$

CROSS-SECTION

SENSITIVE TO BOTH 2-90INT DIPOLE WILSON LINE CORRELATOR & NIVEL 4-POINT QUADRUPOLE CORRELATOR

SOFT BREMSSTRAHLUNG ALSO SHADOWS SUCH CORRELATORS The true of my

JIMWLK RG equations

describe bremsstrakbung of
DIPOLES, AVADRUPOLES, SEXTUPOLES, ...

m an inte hierarchy

Sadly, LLx resummation not good enough for Precision studies

NULX corrections
To RG equations
UNIVERSAL

NLO corrections
to observable
PROCESS
DEPENDENT

*PARTS OF THIS ARE COMPUTED

> NLO JIMWLK "available"

NLO BK MOVE ACCESSIBLE

*PARTS IMPOSED: Eg. RUNNING

COUPLING

$2 \rightarrow N + 2$ amplitudes in the Regge limit of gravity: the Lipatov-ACV approach and the double copy

HIGH-ENERGY SCATTERING IN QCD AND IN QUANTUM GRAVITY AND TWO-DIMENSIONAL FIELD THEORIES

L.N. LIPATOV*

We construct effective actions describing high-energy processes in QCD and in quantum gravity with intermediate particles (gluons and gravitons) having the multi-Regge kinematics. The S-matrix for these effective scalar field models contains the results of the leading logarithmic approximation and is unitary. It can be expressed in terms of correlation functions for two field theories acting in longitudinal and transverse two-dimensional subspaces.

NPB 364 (1991) 614; 157 cites in INSPIRE

Effective action and all-order gravitational eikonal at planckian energies

AMATI, CIAFALONI, VENEZIANO NPB403 (1993)707

Building on previous work by us and by Lipatov, we present an effective action approach to the resummation of all semiclassical (i.e. $O(\hbar^{-1})$) contributions to the scattering phase arising in high-energy gravitational collisions. By using an infrared-safe expression for Lipatov's effective action, we derive an eikonal form of the scattering matrix and check that the superstring amplitude result is reproduced at first order in the expansion parameter R^2/b^2 , where R, b are the gravitational radius and the impact parameter, respectively. If rescattering of produced gravitons is neglected, the longitudinal coordinate dependence can be explicitly factored out and exhibits the characteristics of a shock-wave metric while the transverse dynamics is described by a reduced two-dimensional effective action. Singular behaviours in the latter, signalling black hole formation, can be looked for.

The World as a Hologram

LEONARD SUSSKIND

Wee partons, by contrast, are not subject to Lorentz contraction. This implies that in the Feynman Bjorken model, the halo of wee partons eternally "floats" above the horizon at a distance of order $10^{-13}cm$ as it transversley spreads. The remaining valence partons carry the various currents which contract onto the horizon as in the Einstein Lorentz case.

By contrast, both the holographic theory and string theory require all partons to be wee. No Lorentz contraction takes place and the entire structure of the string floats on the stretched horizon. I have explained in previous articles how this behavior prevents the accumulation of arbitrarily large quantities of information near the horizon of a black hole. Thus we are led full circle back to Bekenstein's principle that black holes bound the entropy of a region of space to be proportional to its area.

J.Math.Phys. 36 (1995) 6377; 3048 cites

In Acknowledgements:

Finally I benefitted from discussions with Kenneth Wilson and Robert Perry, about boosts and renormalization fixed points in light front quantum mechanics and Lev Lipatov about high energy scattering.

These works do not explicitly discuss parton saturation which leads to a perspective I will discuss shortly

$2 \rightarrow N + 2$ amplitudes in the Regge limit of gravity: the Lipatov-ACV approach and the double copy

Double copy between QCD and Gravity amplitudes

Old idea (Kawai-Lewellyn-Tye) based on relations between closed and open string amplitudes – in "low energy" limit between Einstein & Yang-Mills amplitudes

$$M_4^{\text{tree}}(1,2,3,4) = \left(\frac{\kappa}{2}\right)^2 s \, A_4^{\text{tree}}(1,2,3,4) \, A_4^{\text{tree}}(1,2,4,3) \qquad \kappa = 32 \, \pi^2 \, G_N$$

A remarkable "BCJ" color-kinematics duality

Bern, Carrasco, Johansson, arXiv:0805.3993

Tree level $gg \rightarrow gg$ amplitudes (with on shell legs) can be written as

$$i\mathcal{A}_4^{\rm tree} = g^2 \left(\frac{n_s c_s}{s} + \frac{n_t c_t}{t} + \frac{n_u c_u}{u}\right) \quad \text{with the s channel color factor} \quad c_s = -2f^{a_1 a_2 b} f^{ba_3 a_4} \\ \text{kinematic factor} \quad n_s = -\frac{1}{2} \Big\{ \left[(\epsilon_1.\epsilon_2) p_1^\mu + 2(\epsilon_1.p_2) \epsilon_2^\mu - (1 \leftrightarrow 2) \right] \left[(\epsilon_3.\epsilon_4) p_3^\mu + 2(\epsilon_3.p_4) \epsilon_4^\mu - (3 \leftrightarrow 4) \right] \\ \quad + s \left[(\epsilon_1.\epsilon_3) (\epsilon_2.\epsilon_4) - (\epsilon_1.\epsilon_4) (\epsilon_2.\epsilon_3) \right] \Big\}$$

Tree level gravity amplitude obtained fron replacing color factors by kinematic factors

$$i\mathcal{A}_4^{\text{tree}}|_{c_i\to n_i,g\to\kappa/2}=i\mathcal{M}_4^{\text{tree}}=\left(\frac{\kappa}{2}\right)^2\left(\frac{n_s^2}{s}+\frac{n_t^2}{t}+\frac{n_u^2}{u}\right) \quad \text{Significant on-going work on extension to loop amplitudes}$$
 Review: Bern et al., arXiv: 1909.01358

High energy scattering of Gravitons

Same multi-kegge kinematics as previously

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Gravity-QCD double copy in the Regge limit

Lipatov. PLB 116B (1982) 411 Spinor helicity derivation, Liu, arXiv:1811.11710

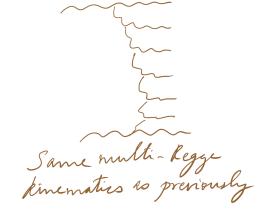
$$\mathcal{M}_n \simeq -s^2 \mathcal{C}(2;3) \frac{-1}{|q_4^{\perp}|^2} \mathcal{V}(q_4;4;q_5) \cdots \frac{-1}{|q_{n-1}^{\perp}|^2} \mathcal{V}(q_{n-1};n-1;q_n) \frac{-1}{|q_n^{\perp}|^2} \mathcal{C}(1;n)$$

Gravitational effective vertex: $V(q_i,i,q_{i+1}) = \Gamma_{i,\mu\nu}(q_i,q_{i+1})\,\epsilon_i^{\mu\nu}(k_i)$ is double copy $\Gamma_i^{\mu\nu}(q_i,q_{i+1}) \equiv 2(C_i^\mu C_i^\nu - N_i^\mu N_i^\nu)$

The C(2;3) and C(1;n) are gravitational "impact factors" that can also be expressed as a double copy of Lipatov's gluon-gluon-reggeized gluon vertex

Writing it as a color kinematics duality is more tricky but can be done

High energy scattering of Gravitons



QED Bremsstrahlung vertex

Lipatov Vertex

Johansson, Sabio Vera,, Campillo, Vasquez-Mozo, arXiv:1310.1680

Classicalization and unitarization in Gravity: The Black Hole N Portrait

The double copy between QCD and gravity amplitudes in the Regge limit suggests that $2 \rightarrow N+2$ grow much faster than QCD amplitudes

When $N=\frac{1}{\alpha_{gr}}$, the dense graviton state unitarized by forming a Black Hole

The Black Hole N Portrait suggests gravity is UV self-complete due classicalization at a weak coupling scale

In the Black Hole N Portrait, Dvali, Gomez, arXiv:1112.3359
Dvali, Guidice, Gomez, Kehagis, arXiv:1010.1415

Gravitons have wavelengths $\lambda = R_S = \sqrt{N} L_P$ $\Delta g_V = \frac{L_P^2}{\lambda^2} = \frac{1}{N} \stackrel{N \to \infty}{\to} 0$

Black Hole evaporation Rate $\int_{H}^{\infty} \int_{H}^{\infty} \int_{H}^{\infty} \int_{R_{s}}^{\infty} \int_{R_{s}}^{$

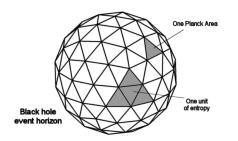
Rs = 26M Schwarzchild radius Lp = ThG Planck Length Mp = Th/G Planck Mass

 $\frac{dM}{dt} = -T^{2}$ and half-life $T = \frac{\hbar^{2}}{T^{3}G} = N^{3/2} L_{P}$

Classicalization and unitarization in Gravity: The Black Hole N Portrait

Bekenstein-Hawking bound

(for a nice discussion, see Bousso, arxlv:1810.01880)



The Bekenstein entropy bound is given by $S < 2\pi ER/\hbar$

In the Black Hole N portrait, this is given by $S \leq 2\pi N Q_S R_S$

which is saturated when
$$N = \frac{1}{\alpha_{gr}} \rightarrow S_{Bek} = \frac{1}{\alpha_{gr}}$$

Here E = N Q_s is the energy in a critically packed volume = R_s^3 of quanta, and $Q_s = 1/R_s$

From our previous discussion,
$$S_{Bek} = \frac{1}{\alpha_{gr}} = \frac{R_S^2}{L_P^2} = \frac{Area}{4G} = S_{BH}$$

The entropy can also be expressed in terms of a Goldstone decay constant corresponding to spontaneous breaking of Poincare invariance by the graviton condensate: $f^2 = N Q_s^2$ Dvali. arXiv:0907.07332 The entropy in these units is S_{BH} = Area \times f^2

At the critical point $N=\frac{1}{\alpha_{gr}}$, the saturation of the BH bound also leads to unitarization!

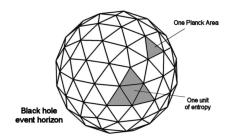
$$G_{2\rightarrow N}^{3HNP} = N / dy e^{S_{BH}} \simeq Me^{-N} / Ne^{-N} = e^{-\frac{1}{dy} + \frac{1}{dy}} \sim O(1)$$
Dvali, Gomez, Isermann, Lust, Stieberger arXiv:1409.7405

Away from the critical point, the cross-section for BH formation drops exponentially!

Classicalization and unitarization in Gravity: The Black Hole N Portrait

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(for a nice discussion, see Bousso, arxlv:1810.01880)



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How to connect to Lipatov-ACV amplitude language? This entropy factor arises from cancellation of infrared divergences In the Sudakov form factor with real soft contributions – to give a finite result a la BFKL

Addazi, Bianchi, Veneziano, ar Xiv:1611.03643

The CGC/BHNP correspondence

Dvali, RV, in preparation

From the previous discussion, we begin to glimpse the elements of the CGC/BHNP correspondence with a region of space $R_S = Q_S^{-1}$ of the CGC mapped on to a Black Hole on analogous radius determined by the saturation of the graviton condensate:

The scale at which this occurs is a weak coupling scale in both theories:

for BHNP:
$$Q_S^{-1} = \sqrt{N} M_P^{-1}$$
,
for CGC: $Q_S = \sqrt{N} \Lambda_{\rm QCD}$

This scale is determined by a maximal occupancy of $N=\frac{1}{\alpha}$ which is where (perturbative) unitarization occurs Its energy-dependent: the "wee-er" the parton, the quicker it saturates...

The direct correspondence between the two theories only occurs at $Q=Q_S$ where the physics is universal - It is independent of the details of the dynamics of the wee partons

The existence of a double copy between the two theories is however suggestive that it can be extended to a "classical" double copy between the gluon and graviton classical shockwaves

Review on classical double copy: C. White, arXiv:1708.07056

The CGC/BHNP correspondence: Bekenstein entropy

Dvali, RV, in preparation

We can use this direct correspondence to conjecture the maximal Bekenstein entropy of the CGC

Since the typical momenta of saturated wee gluons is Q_S , the CGC energy is N Q_S in a "color domain" of radius Q_S^{-1} - and the entropy argument goes through as previously:

$$S_{CGC} = \frac{1}{\alpha_S}$$

Inside a confining radius of Λ_{QCD}^{-1} , of the Lorentz contracted proton at high energies, there can be $\frac{Q_S^2}{\Lambda_{QCD}^2}$ = N domains Hence, $S_{proton} = N S_{CGC} = \frac{1}{\alpha_S^2}$

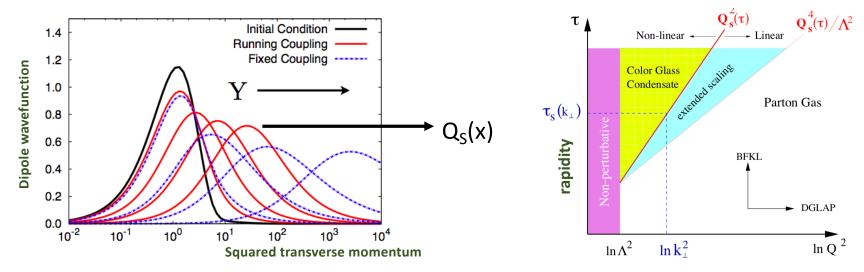
Further, since one can have 1/ α_S wee partons in a rapidity interval Y, this gives $\frac{dS_{proton}}{dY} = \frac{1}{\alpha_S}$

Other discussions in the literature:
Kovner, Lublinsky, Serino, arXiv:1806.01089
Armesto, Dominguez, Kovner, Lublinsky, Skokov, arXiv:1901.08080

The CGC/BHNP correspondence: Bekenstein entropy

Dvali, RV, in preparation

One can also understand the CGC entropy in the language of a Goldstone decay constant characterizing the breaking of the Poincare symmetry of partons due to the formation of the "soliton-like" CGC state



This "UV" scale characterizes the onset of screening (shadowing) and is given by $f^2 = \frac{Q_S^4}{\Lambda_{QCD}^2}$

Thus, just as in the BHNP, the CGC entropy satisfies an area law:

$$S_{CGC}$$
 = Area * f² , where Area = $4\pi R_S^2$

Final remarks

The CGC/Black Hole correspondence we have outlined also extends to the evolution of the CGC/Glasma and Black Hole evaporation via Hawking radiation – the analogy is however not valid after a "quantum breaking" time beyond which the CGC forms a QGP

The study of "soft factors" in the CGC and the relation to "cusp anomalous dimensions" is under investigation in a worldline QFT formalism of radiation from sem—classical worldlines

Feal, Tarasov, RV, in progress

The classical double copy of CGC and saturated graviton states is also under investigation and can provide more quantitative evidence of the correspondence outlined

Raj, RV, in progress