

Massive neutron stars with quarks

Vector–interaction–enhanced bag model

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The Dyson–Schwinger equation formalism

In-medium quark propagator

$$S(p^2, \tilde{p}_4)^{-1} = i\vec{\gamma}\vec{p}A(p^2, \tilde{p}_4) + i\gamma_4\tilde{p}_4C(p^2, \tilde{p}_4) + B(p^2, \tilde{p}_4),$$

The Dyson–Schwinger equation for the quark propagator

$$S(p^2, \tilde{p}_4)^{-1} = i\vec{\gamma}\vec{p} + i\gamma_4\tilde{p}_4 + m + \Sigma(p^2, \tilde{p}_4),$$

with $\tilde{p}_4 = p_4 + i\mu$, and the self-energy term

$$\Sigma(p^2, \tilde{p}_4) = \int \frac{d^4q}{(2\pi)^4} g^2(\mu) D_{\rho\sigma}(p-q, \mu) \frac{\lambda_\alpha}{2} \gamma_\rho S(q^2, \tilde{q}_4) \Gamma_\sigma^\alpha(q, p, \mu),$$

Truncations:

$$g^2 D_{\rho\sigma}(p-q) = \frac{1}{m_G^2} \theta(\Lambda^2 - \vec{q}^2) \delta^{\rho\sigma},$$

$$\Gamma_\sigma^\alpha(q, p, \mu) = \frac{\lambda^\alpha}{2} \gamma_\sigma.$$

Derived gap equations:

$$\tilde{p}_4^2 C(p^2, \tilde{p}_4) = \tilde{p}_4^2 + \frac{8}{3m_G^2} \int \frac{d^4q}{(2\pi)^4} \frac{\tilde{p}_4 \tilde{q}_4 C(q^2, \tilde{q}_4)}{\vec{q}^2 A^2(q^2, \tilde{q}_4) + \tilde{q}_4^2 C^2(q^2, \tilde{q}_4) + B^2(q^2, \tilde{q}_4)},$$

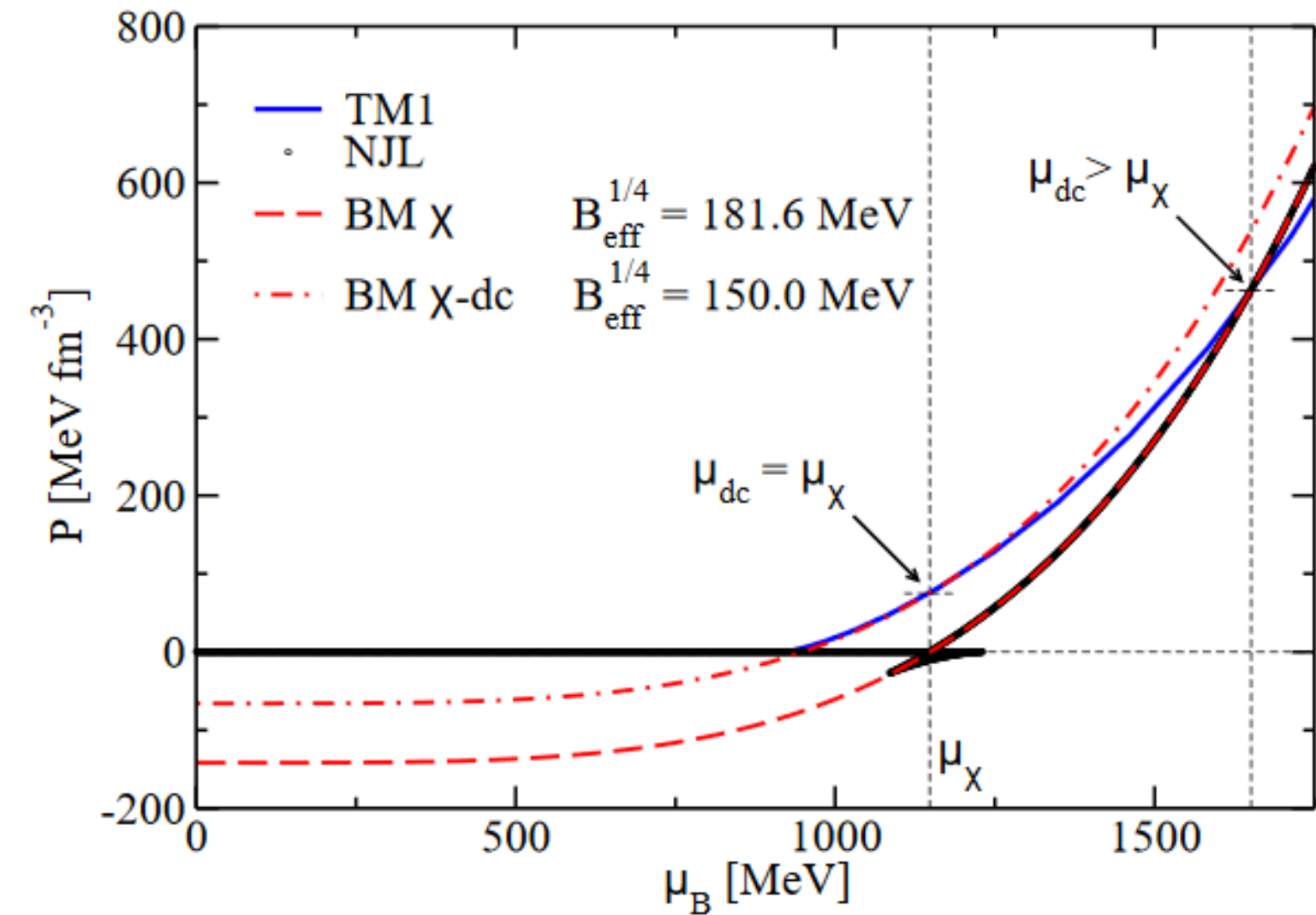
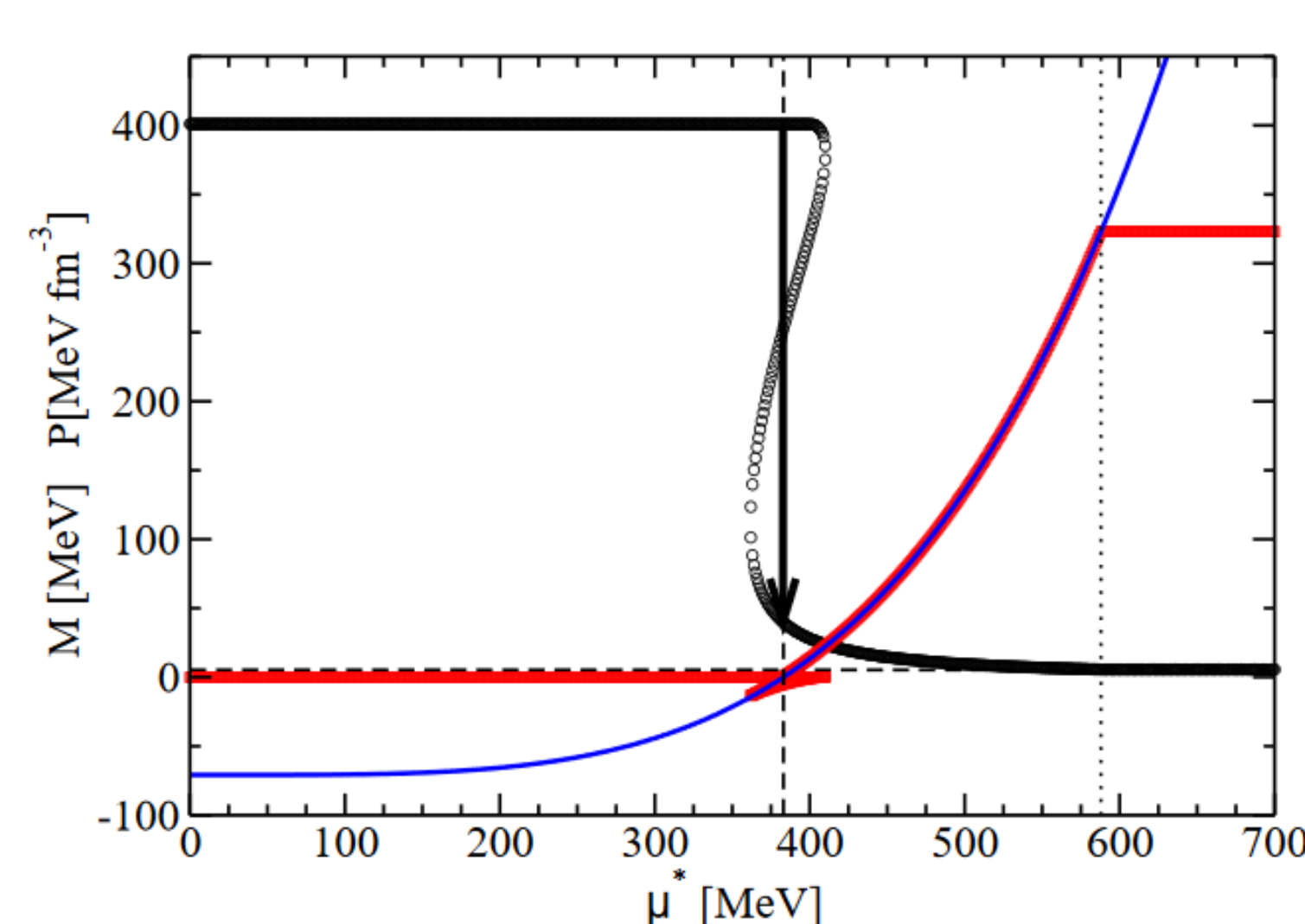
$$B(p^2, \tilde{p}_4) = m + \frac{16}{3m_G^2} \int \frac{d^4q}{(2\pi)^4} \frac{B(q^2, \tilde{q}_4)}{\vec{q}^2 A^2(q^2, \tilde{q}_4) + \tilde{q}_4^2 C^2(q^2, \tilde{q}_4) + B^2(q^2, \tilde{q}_4)}$$

The DSE–derived single flavor pressure (left picture, red line or right picture, black line)

$$P_{FG} = Tr Ln S^{-1} = 2N_c \int_\Lambda \frac{d^4q}{(2\pi)^4} Ln(\vec{p}^2 + \tilde{p}_4^2 + B^2),$$

and

$$P_l = -\frac{1}{2} Tr \Sigma S = \frac{3}{4} m_G^2 (\mu - \mu^*)^2 - \frac{3}{8} m_G^2 (B - m)^2.$$



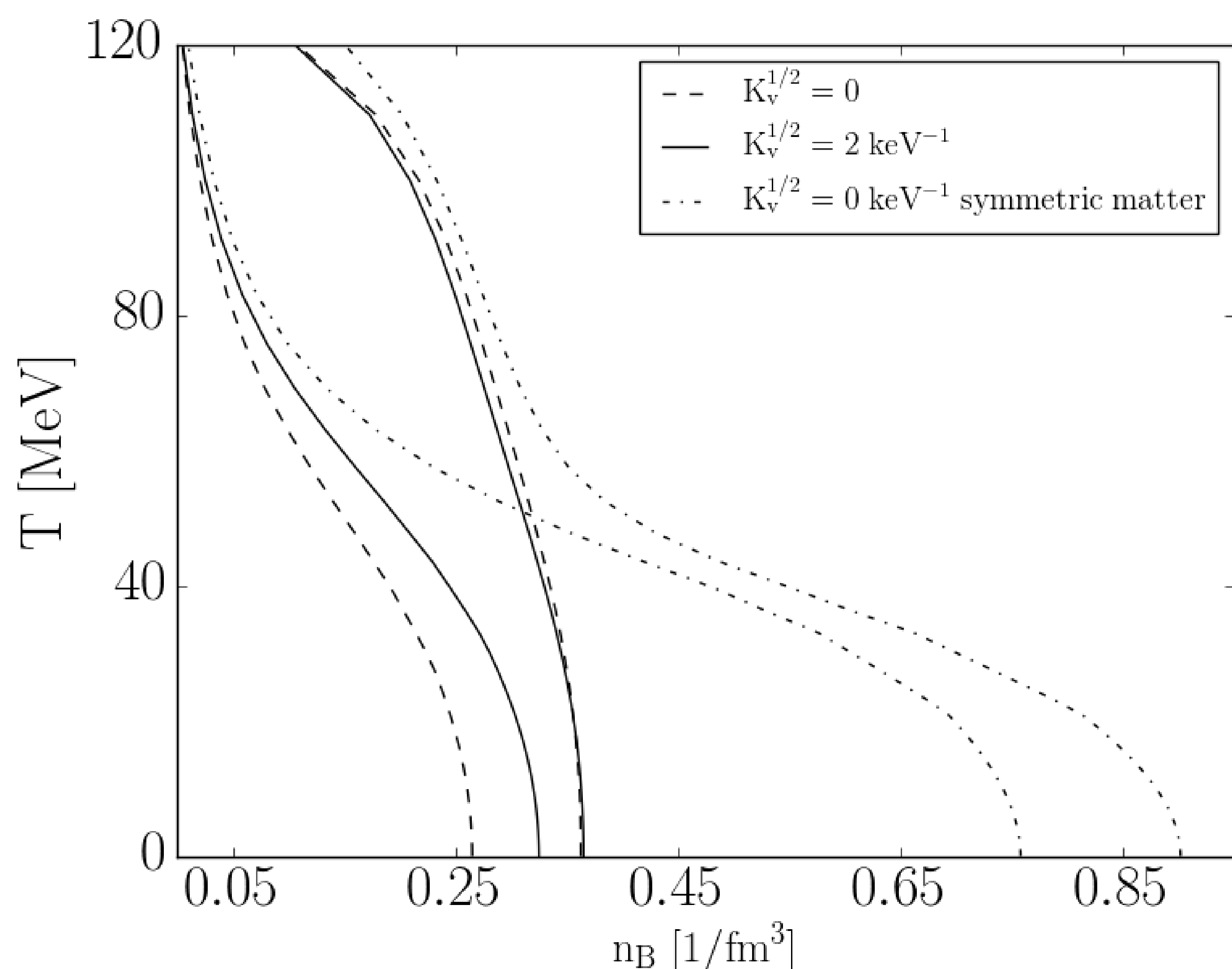
$$P_f(\mu_f) = P_{FG,f}^{kin}(\mu_f^*) - B_{\chi,f}.$$

$$\mu_f = \mu_f^* + K_v n_{FG,f}(\mu^*)$$

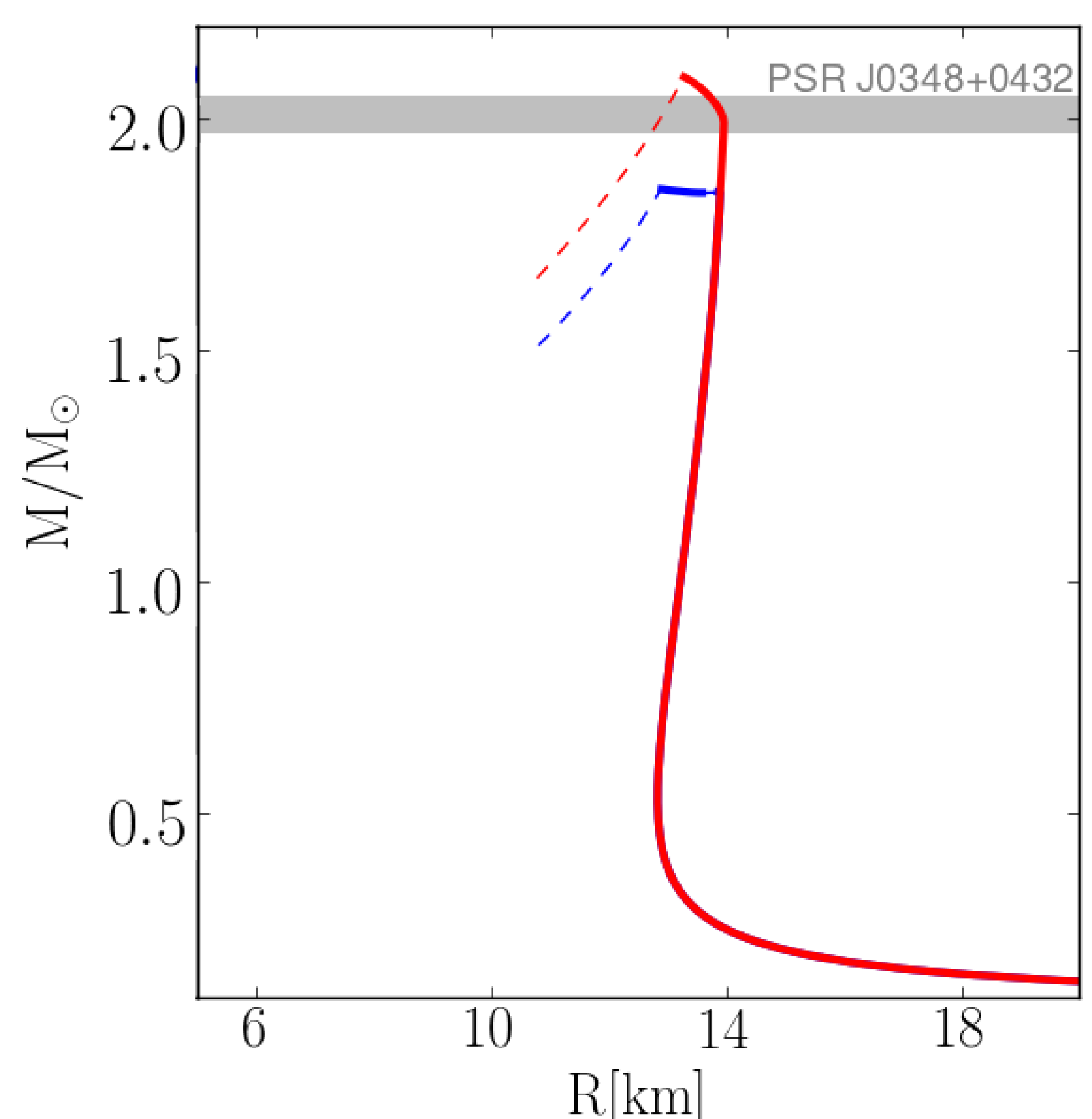
$$P^Q = \sum_{\mu_B} P_f(\mu_f) + B_{dc}$$

$$\mu_B = \mu_u + 2\mu_d$$

Phase diagram



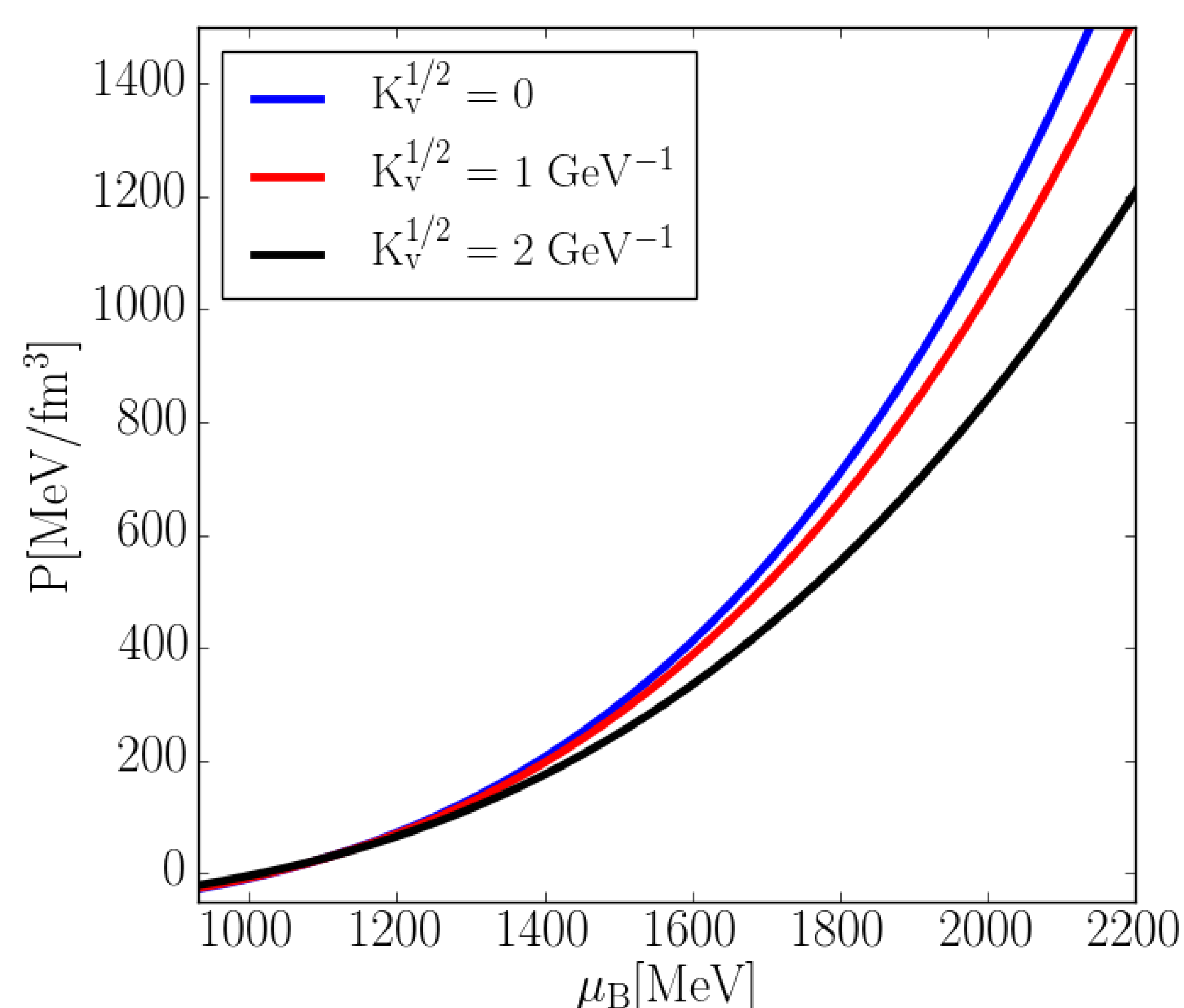
Neutron star mass–radius relation



Repulsive vector interaction essential for the presence of massive hybrid neutron stars is realised by redefining the pressure as

$$P_f(\mu_f) = P_{FG,f}^{kin}(\mu_f^*) + \frac{K_v}{2} n_{FG,f}^2(\mu_f^*) - B_{\chi,f}.$$

The impact on the single flavor EoS:



vBag equation of state

Single flavor chemical potential

$$\mu_f = \mu_f^* + K_v n_{FG,f}(\mu^*)$$

Baryochemical potential

$$\mu_B = \mu_u + 2\mu_d$$

Charge chemical potential

$$\mu_C = \mu_u - \mu_d$$

Particle number density

$$n_f(T, \mu_f) = n_{FG,f}(T, \mu_f^*)$$

Baryon density

$$n_B = \sum_f \frac{1}{3} n_f(T, \mu_f^*)$$

Single flavor pressure

$$P_f(T, \mu_f) = P_{FG,f}^{kin}(T, \mu_f^*) + \frac{K_v}{2} n_{FG,f}^2(\mu_f^*) - B_{\chi,f}$$

Total pressure

$$P^Q = \sum_f P_f(T, \mu_f) + B_{dc}(T, \mu_C)$$

Single flavor energy density

$$\epsilon_f(T, \mu_f) = \epsilon_{FG,f}^{kin}(T, \mu_f^*) + \frac{K_v}{2} n_{FG,f}^2(\mu_f^*) + B_{\chi,f}$$

Total energy density

$$\epsilon^Q = \sum_f \epsilon_f(T, \mu_f^*) - B_{dc}(T, \mu_C) + T \frac{\partial B_{dc}(T, \mu_C)}{\partial T} + \mu_C \frac{\partial B_{dc}(T, \mu_C)}{\partial \mu_C}$$

Single flavor entropy

$$s_f(T, \mu_f) = s_{FG,f}^{kin}(T, \mu_f^*)$$

Total entropy

$$s^Q = \sum_f s_f(T, \mu_f) + \frac{\partial B_{dc}(T)}{\partial T}$$



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