# **GRETINA: Status and Recent Results**

- GRETINA
  - Device Overview
  - Enhancements
  - Physics Campaigns
- Results
  - Selected highlights
- Future
  - GRETINA
  - GRETA

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### Gamma-Ray Energy Tracking Array

#### **GRETA concept for a shell of closely packed Ge crystals**

-0.05

-0.15

-0.2

-0.4

Pulse-shape analysis

- Combines highly segmented, hyper-pure germanium crystals with advanced digital signal processing techniques
- Identify the position and energy of γ-ray interaction points within a compact "shell" of detectors
- Track γ-ray path both within and between detector elements, using the angle-energy relation of the Compton scattering process



- Efficiency, Energy Resolution, Peak-to-Total
- Also, Doppler reconstruction! (RIBs inverse kinematics)

NUSPIN 2017 Workshop

 $\theta_3, E_3$ 

 $\frac{1}{200} \frac{1}{400} \frac{1}{0} \frac{1}{0} \frac{1}{200} \frac{1}{400} \frac{1}{100} \frac{1}{100$ 

## GRETINA



# Gamma Ray Energy Tracking In beam Nuclear Array



**\$20M** Funded by US- DOE Nuclear Physics Office











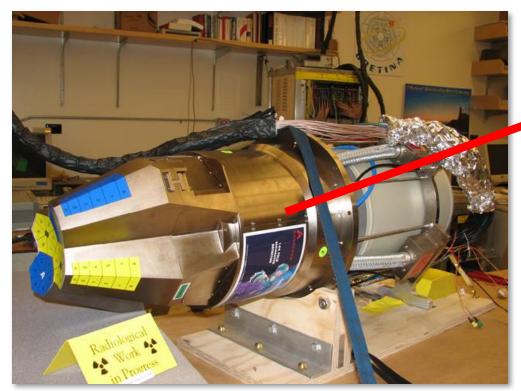


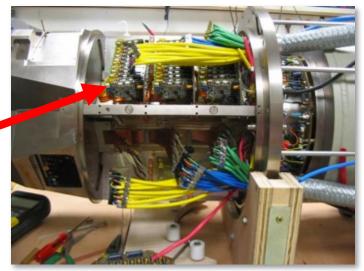
- A first realization of a Tracking Array
   Optimized for fast beam experiments
- Coverage ~  $\frac{1}{4}$  of  $4\pi$  solid angle
- 28 36-fold segmented Ge crystals
  - 7 Quad Modules
- Mechanical support structure
- Data acquisition system
- Data processing software



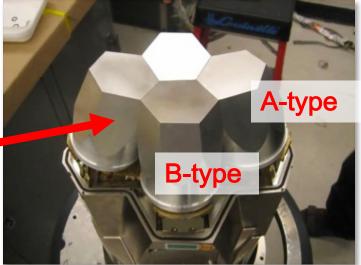
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## **Detector Quad Modules**





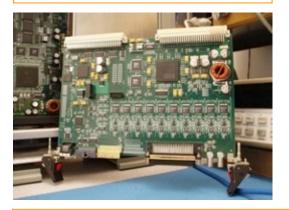
36 segments/crystal
4 crystal/ module
148 signal channels /module
Cores Cold FETs
Segments Warm FETs





# **Electronics and DAQ System**

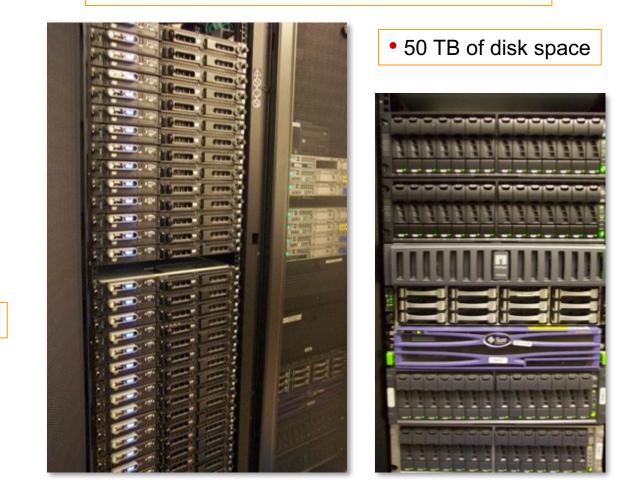
Digitizer (LBNL) 14bit, 100 MHz CC Energy: 2.5, 5, 10, 30 MeV Leading edge time Pulse shape



#### Trigger Timing & Control (ANL)



• Installed 62 nodes, 2 CPU/node, 4 core/ CPU



#### 56 Nodes $\geq$ 20000 gammas/s

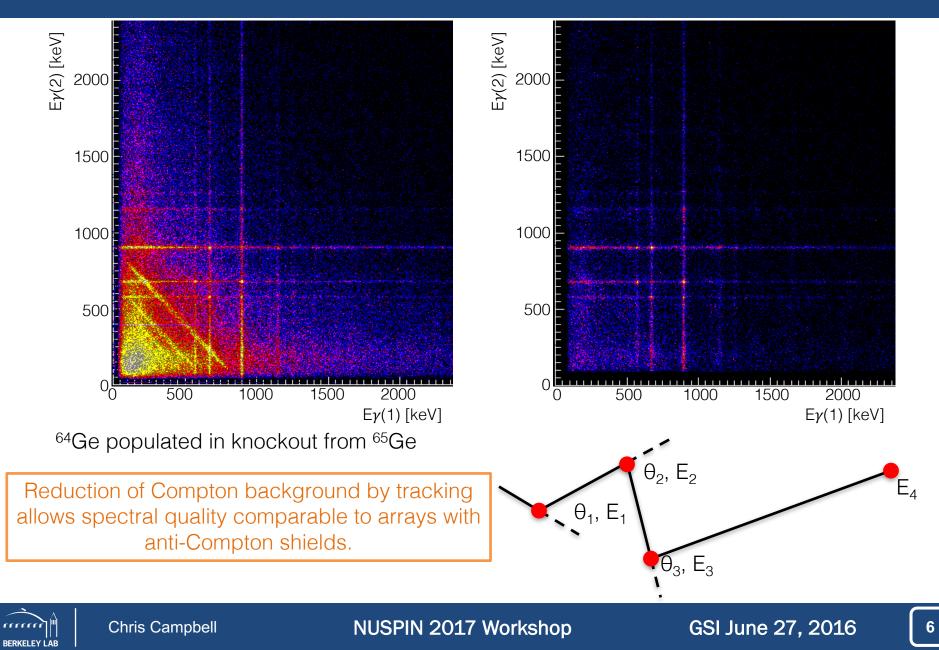


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GSI June 27, 2016

# Tracking: Compton Rejection



# GRETINA "Enhancements"

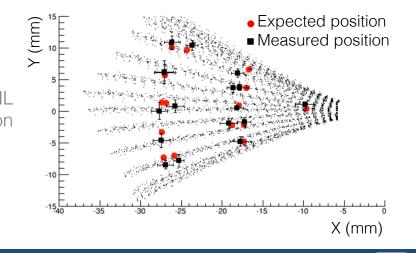
GRETINA enhancements have been supported by the U.S DOE-NP Office since completion of the project with 7 quad modules in 2011.



- Growth of the array through continued procurement of quad modules (and associated infrastructure) at a rate of ~ 1 per year
  - Purchase of Q8 Q12
  - Electronics (digitizers, trigger modules, cables and power for each module)
  - Supplementation of the computing cluster to support the increased data rates of the larger array

#### • Firmware development and modifications

- Improved DAQ stability
- High-rate energy resolution performance improvements
- Refined timing performance
- Double throughput? (50MHz trace)
- Signal decomposition and tracking
  - In-depth coincidence scanning of Q4 at LBNL
  - Extended capabilities in signal decomposition algorithms
  - Enhanced simulation including the effects of signal decomposition





### Physics campaigns: Auxiliary devices

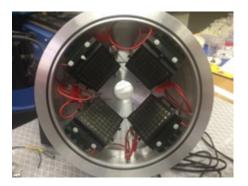
#### S800

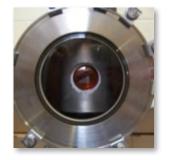




#### TRIPLEX



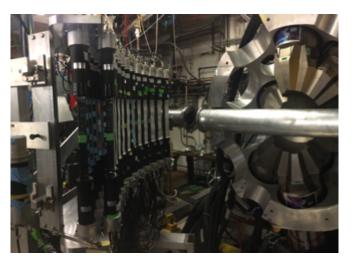




LH Target



#### CHICO II

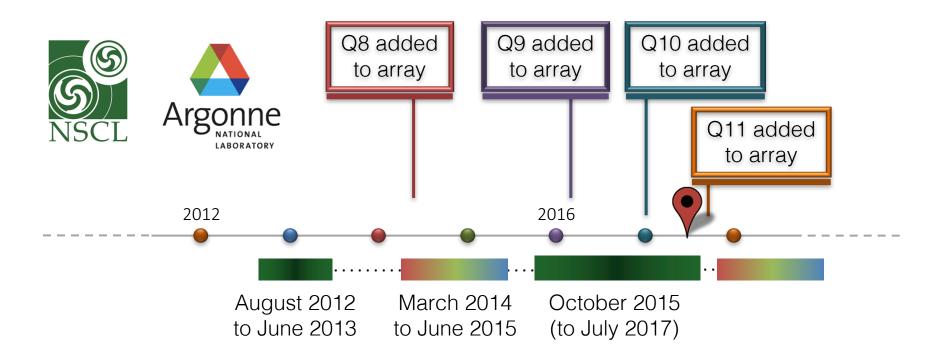


LENDA



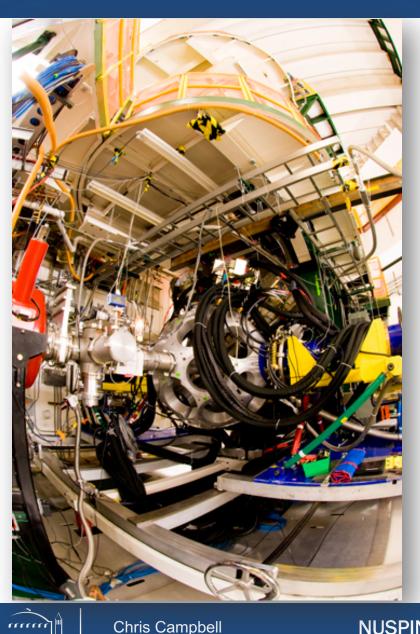
**Chris Campbell** 

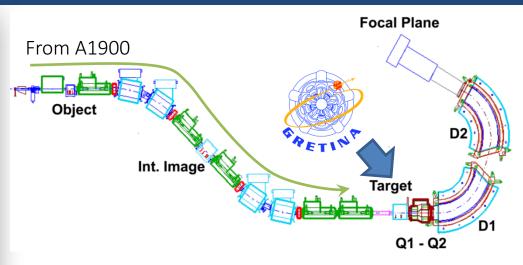
# **GRETINA Physics Campaigns: Overview**



- First physics campaigns at NSCL and ANL (ATLAS/CARIBU) have already produced more than 25 physics papers
- Second NSCL campaign is ongoing, with 24 PAC approved experiments (~3600 hours)

# First NSCL Physics Campaign (2012-2013)





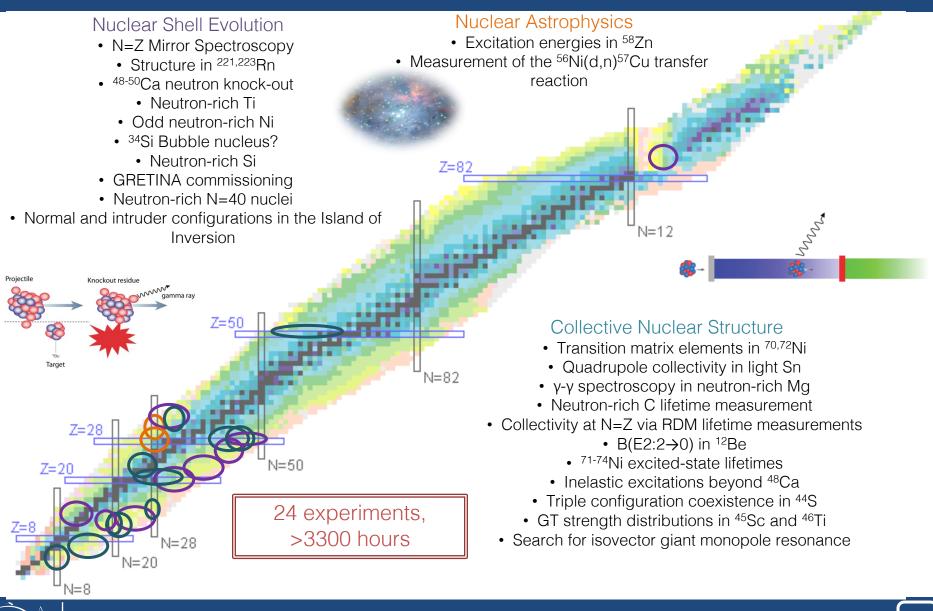
- First physics campaign coupled GRETINA with the S800 spectrograph for fast fragmentation beam (in-beam) spectroscopy
- Reaction channel selection is possible event-by-event in the S800

#### Challenges:

- High beam velocities ( $\beta \approx 0.35$ ) 2mm spatial resolution of GRETINA optimizes Doppler correction
- Low beam rates (as low as a few pps) efficiency of GRETINA, in singles and γγ is critical



# First NSCL Physics Campaign (2012-2013)



### First ANL Physics Campaign (2014-2015)

#### **18** experiments : 3350 hours



CARIBU and Stable beams

**Position Resolution** 

Good response for high-energy gammas

Polarization sensitivity



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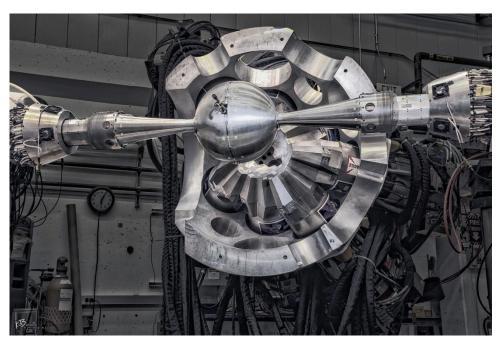
# First ANL Physics Campaign (2014-2015)

N=126

#### CARIBU Experiments

- Shape coexistence in neutron-rich Ge isotopes <sup>72</sup>Ge and <sup>76</sup>Ge
  - Octupole strength in <sup>225</sup>Ra
- Octupole strength in neutron-rich nuclei near A~140
  - Shape coexistence in <sup>100</sup>Zr
  - Role of the  $g_{9/2}$  orbital in <sup>71</sup>Zn

First physics campaign at ANL made use of the unique opportunities with CARIBU beams, and the combination of GRETINA with particle detectors such as CHICO-2 and the WashU Phoswich Wall.

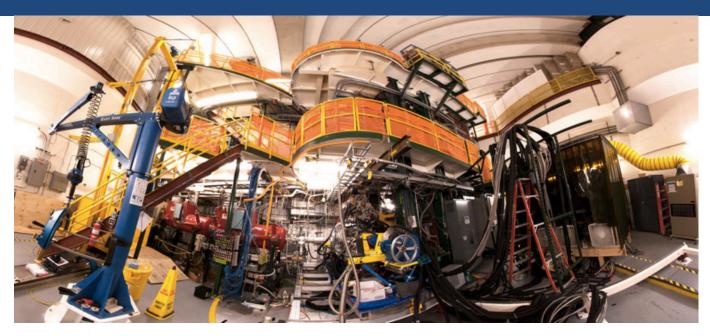


#### ATLAS Stable Beam Experiments

- Comparison of the performance of GRETINA and Gammasphere
  - High-energy gamma-ray performance
  - The polarization sensitivity of GRETINA
  - Multi-particle-hole states in neutron-rich <sup>34</sup>P
- Understanding the low-energy enhancement of the photon-strength function



# Current NSCL Campaign (2015-ongoing)



- Second NSCL campaign ongoing, ends July 2017
- 24 PAC approved experiments (~3600 hours)
- Campaign broadly covers the light to mediummass nuclei and includes experiments with:
  - Recoil-distance lifetime measurements, and lifetimes through back-tracking
  - Gamma-ray + particle + neutron triple coincidences with LENDA + S800
  - Liquid-hydrogen target measurements



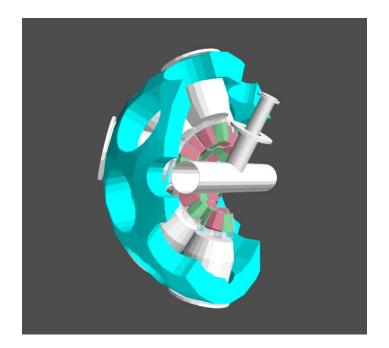


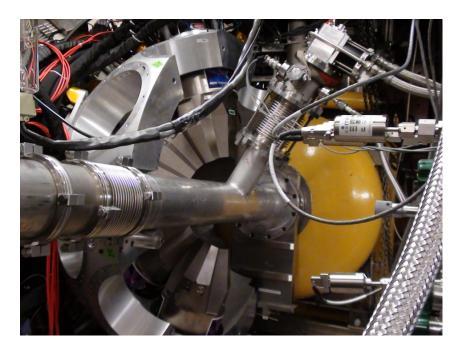
# Now a biased selection of interesting experiments and results



## Thick-target (p,p'), knockout, fragmentation @ NSCL

- inelastic only,  $\gamma$ -ray spectroscopy
- 200 mg/cm<sup>2</sup> target
- 1-100 pps
- simulations crucial for  $\gamma$ -ray efficiencies and target thickness





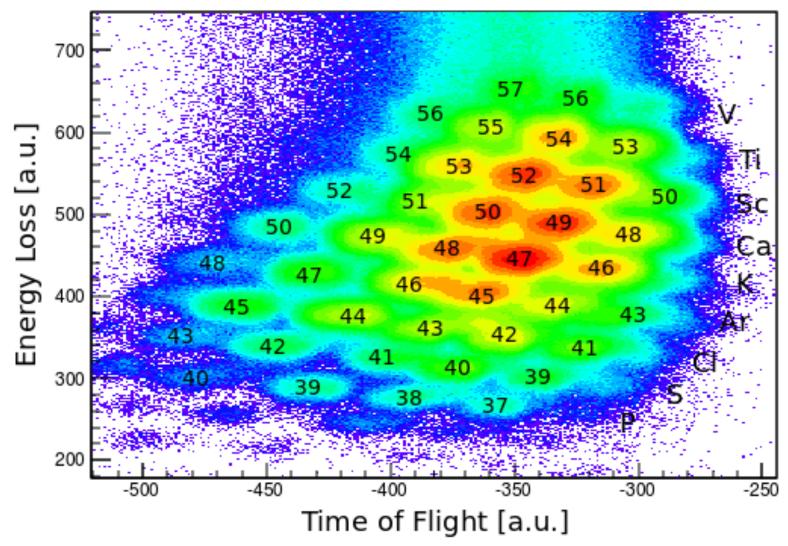
**GRETINA + NSCL/Ursinus LH target + S800** NSCL Experiments 11035, 11037, 12016, 10010



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## NSCL Expt. 11037

Incoming PID





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### Octupole Deformation in <sup>144</sup>Ba

PRL 116, 112503 (2016)

PHYSICAL REVIEW LETTERS

week ending 18 MARCH 2016

#### Direct Evidence of Octupole Deformation in Neutron-Rich <sup>144</sup>Ba

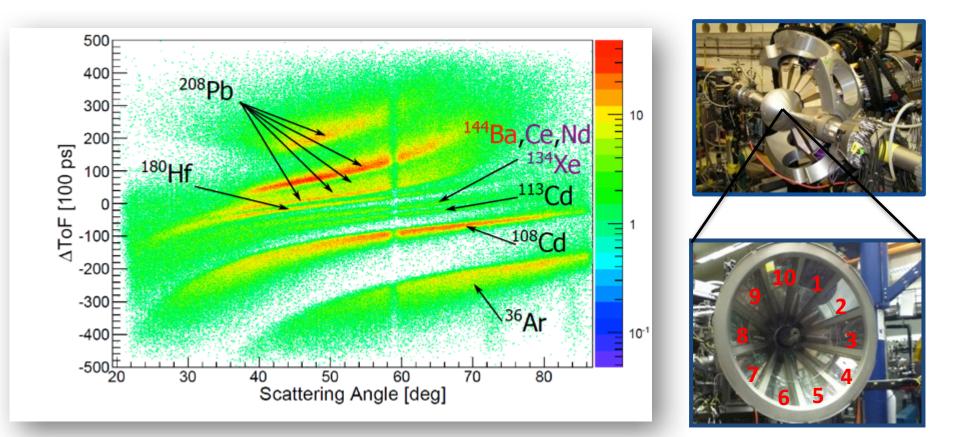
B. Bucher,<sup>1,\*</sup> S. Zhu,<sup>2</sup> C. Y. Wu,<sup>1</sup> R. V. F. Janssens,<sup>2</sup> D. Cline,<sup>3</sup> A. B. Hayes,<sup>3</sup> M. Albers,<sup>2</sup> A. D. Ayangeakaa,<sup>2</sup> P. A. Butler,<sup>4</sup> C. M. Campbell,<sup>5</sup> M. P. Carpenter,<sup>2</sup> C. J. Chiara,<sup>2,6,†</sup> J. A. Clark,<sup>2</sup> H. L. Crawford,<sup>7,‡</sup> M. Cromaz,<sup>5</sup> H. M. David,<sup>2,§</sup> C. Dickerson,<sup>2</sup> E. T. Gregor,<sup>8,9</sup> J. Harker,<sup>2,6</sup> C. R. Hoffman,<sup>2</sup> B. P. Kay,<sup>2</sup> F. G. Kondev,<sup>2</sup> A. Korichi,<sup>2,10</sup> T. Lauritsen,<sup>2</sup> A. O. Macchiavelli,<sup>5</sup> R. C. Pardo,<sup>2</sup> A. Richard,<sup>7</sup> M. A. Riley,<sup>11</sup> G. Savard,<sup>2</sup> M. Scheck,<sup>8,9</sup> D. Seweryniak,<sup>2</sup> M. K. Smith,<sup>12</sup> R. Vondrasek,<sup>2</sup> and A. Wiens<sup>5</sup> <sup>1</sup>Lawrence Livermore National Laboratory, Livermore, California 94550, USA <sup>2</sup>Argonne National Laboratory, Argonne, Illinois 60439, USA <sup>3</sup>University of Rochester, Rochester, New York 14627, USA <sup>4</sup>University of Liverpool, Liverpool L69 7ZE, United Kingdom <sup>5</sup>Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA <sup>6</sup>University of Maryland, College Park, Maryland 20742, USA <sup>7</sup>Ohio University, Athens, Ohio 45701, USA <sup>8</sup>University of the West of Scotland, Paisley PA1 2BE, United Kingdom <sup>9</sup>SUPA, Scottish Universities Physics Alliance, Glasgow G12 8QQ, United Kingdom <sup>10</sup>CSNSM, IN2P3-CNRS, bâtiment 104-108, F-91405 Orsay Campus, France <sup>11</sup>Florida State University, Tallahassee, Florida 32306, USA <sup>12</sup>University of Notre Dame, Notre Dame, Indiana 46556, USA





# CHICO II and GRETINA

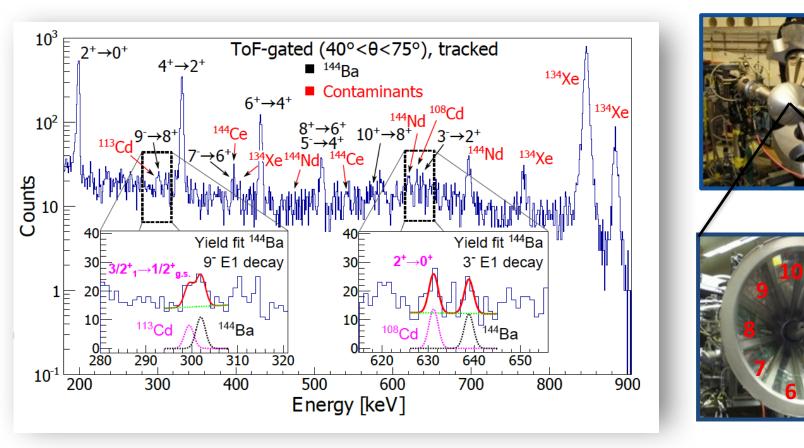
~ 8000 pps 650 MeV <sup>144</sup>Ba Beam on a <sup>208</sup>Pb target (1 mg/cm<sup>2</sup>) for ~ 10 days Important stable contaminants (same m/q): <sup>134</sup>Xe<sup>26+</sup>, <sup>144</sup>Nd<sup>28+</sup>, <sup>180</sup>Hf<sup>35+</sup>, <sup>108</sup>Cd<sup>21+</sup>, <sup>113</sup>Cd<sup>22+</sup>





# CHICO II and GRETINA

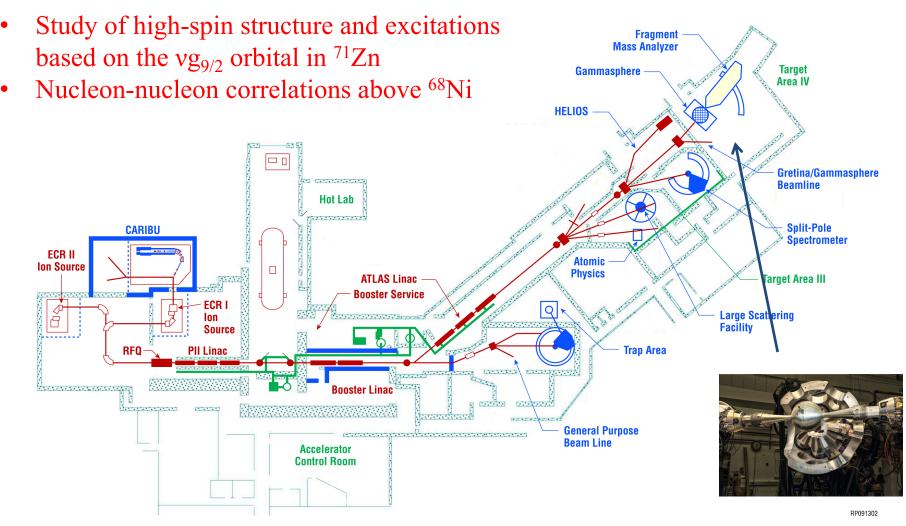
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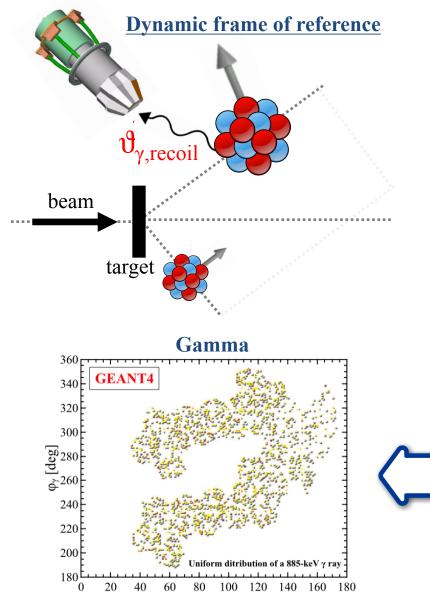
#### The <sup>48</sup>Ca+<sup>70</sup>Zn experiment

• Test of heavy-ion transfer reactions with GRETINA and CHICO2



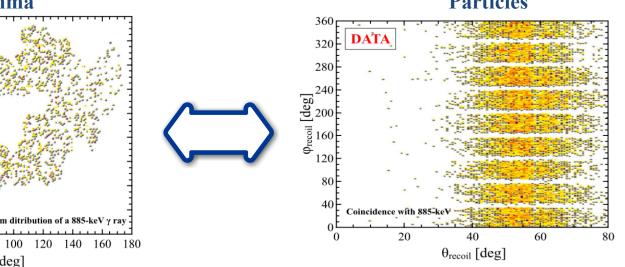
<sup>48</sup>Ca+<sup>70</sup>Zn  $\rightarrow$  <sup>47</sup>Ca+<sup>71</sup>Zn @ 170 MeV (30 % above C.B.)

#### **Angular distributions**



 $\theta_{\gamma}$  [deg]

- Simulation of an isotropic gamma ray 1) (GEANT4)
- 2) Process the gamma ray as real data (tracking)
- 3) **Extract particle angular distribution** gated on that gamma ray
- **4**) Fold gammas and particles event by event



#### **Particles**

## **Experimental Method: Photon Strength Function**

Excitation Energy [MeV]

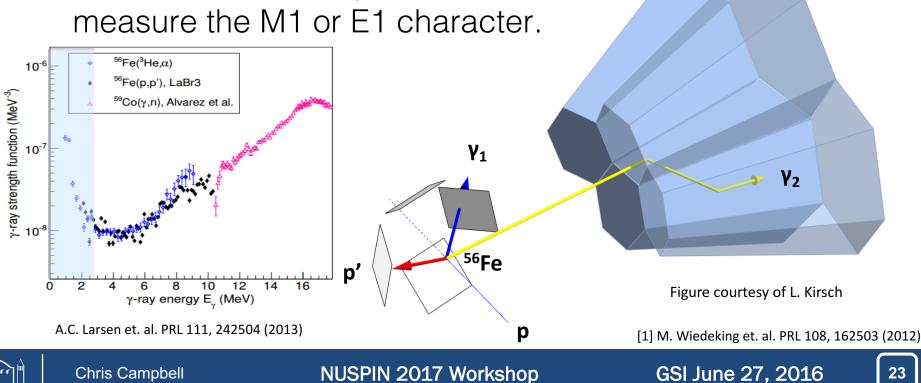
actions Direct

Fusion. evaporation

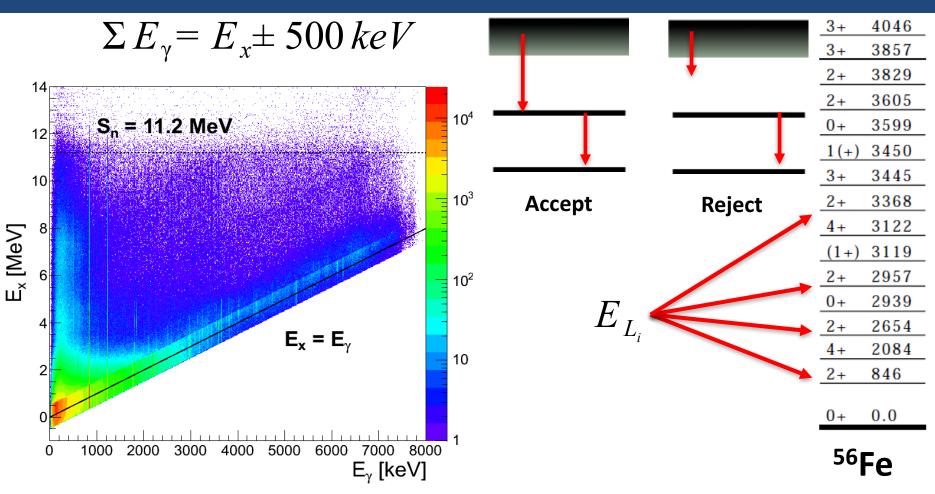
Spin

10

- Use a direct reaction to populate states in quasi-continuum.
- Extract PSF using a new model independent method [1].
- Identify events in the "up-bend"
  - Use GRETINA as a polarimeter to measure the M1 or E1 character.



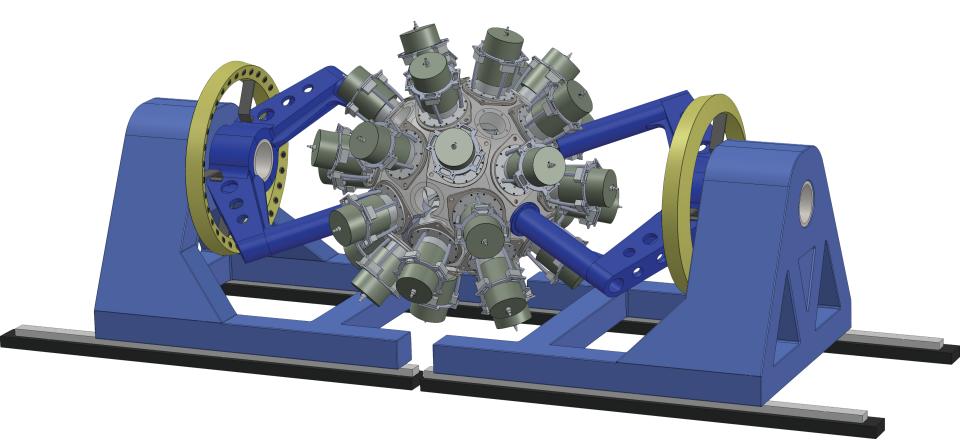
# **Calorimeter Gating**



- Measure intensity of feeding to first four 2<sup>+</sup> states.
- Correct for efficiency, branching, and background.



#### Future: GRETINA to GRETA





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### Future of GRETINA

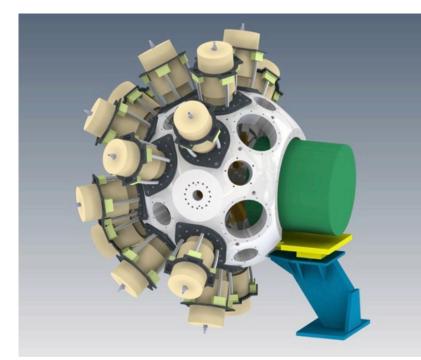
#### **Near Term**

- Move to Argonne for 2<sup>nd</sup> physics campaign, now @FMA
- Improve rate performance
- Next move(s) TBD by community + DOE

#### Long Term

- Serves FRIB for initial operation
- Quad modules will be incorporated into GRETA

#### New FMA entrance quad



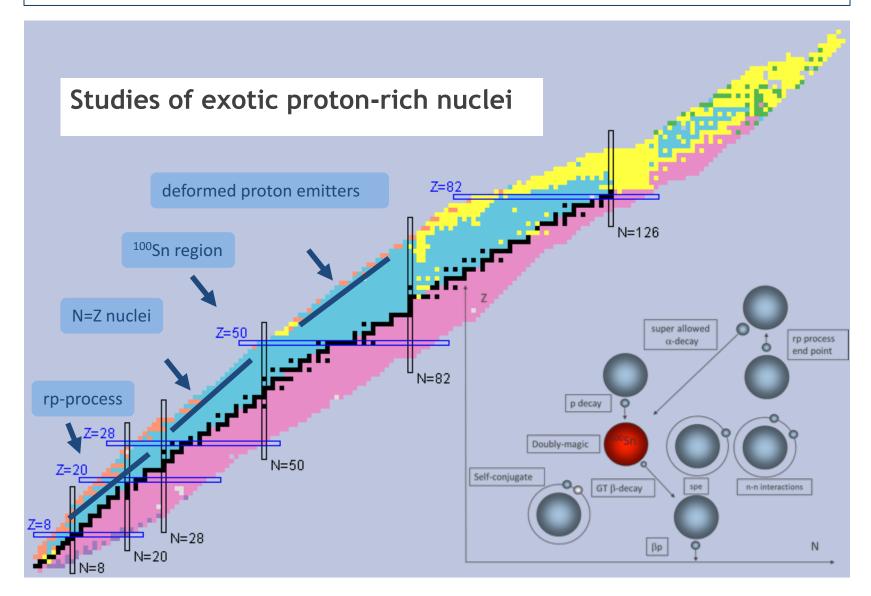
Will fit inside GRETINA frame (most forward detector ring removed)

Larger solid angle 12 msr (8 msr with FMA alone now, 2 msr with GS)

Two hemis at 90 degrees to the beam axis (>10 clusters) high  $\gamma$ -ray efficiency, Doppler correction, polarization



#### • ATLAS 2.0 GRETINA@FMA campaign ~ October 2017 ....



#### GRETA: A premier $\gamma$ -ray tracking detector for FRIB

The Facility for Rare Isotope Beams (FRIB), is currently being constructed at Michigan State University. It will be a world leading accelerator facility to understand the properties of rare nuclear isotopes and how elements are synthesized.

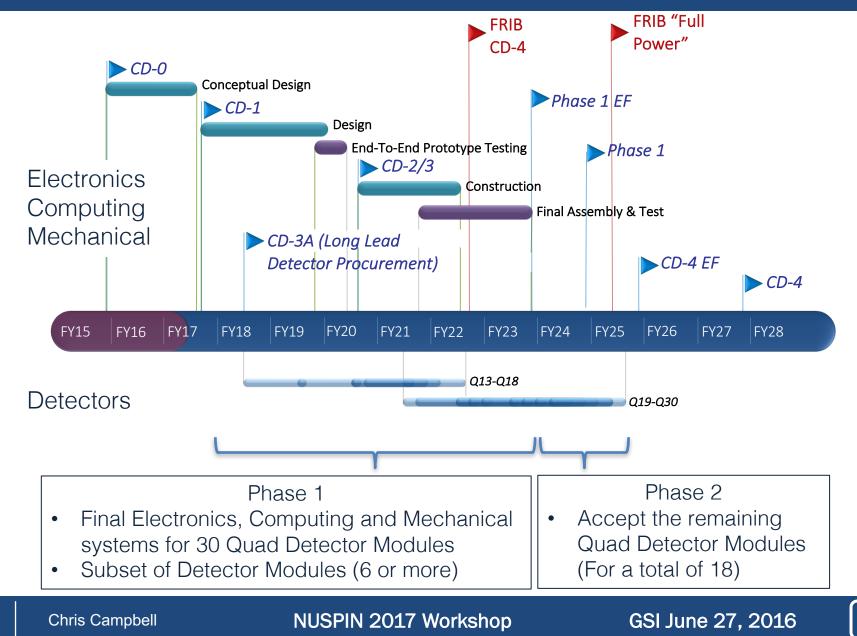
A high resolution, high efficiency γ-ray detector array (GRETA) is an essential instrument for FRIB, required to leverage the full scientific capabilities.







# Phased approach to optimize delivery of a $\gamma\text{-ray}$ tracking array capability for FRIB



BERKELEY LAB

### **GRETINA/GRETA** User Community

Established and Engaged User Community

#### **GRETINA/GRETA Users Executive Committee (GUEC)**

Heather Crawford (chair) Hiro Iwasaki Darek Seweryniak Lew Riley Mark Riley



Over 200 active Users

Coordinated with the FRIB Users Organization

Establishing Working Groups

First workshop planned (August 2<sup>nd</sup> 2017)

<u>http://gretina.lbl.gov/</u>

NUSPIN 2017 Workshop

GSI June 27, 2016

http://greta.lbl.gov/

### Summary

**GRETINA**, a first implementation of a tracking array, is a powerful tool for in-beam gamma-ray spectroscopy.

Physics campaigns were completed successfully at ANL and NSCL Results show very good performance and new and exciting physics NSCL 2.0 : October 2015 – July 2017

Next campaign at ATLAS  $\rightarrow$  FMA and improved CARIBU beams

The path from GRETINA to GRETA has been proposed:

**GRETINA (12 Quads) thru FRIB early operation** 

GRETA Phase 1 aims to take over after ~1yr : (12 + 6 =) 18 Quads + new systems

GRETA Phase 2 aims to expand GRETA+FRIB science reach by adding Quads until Full (30 Quad) GRETA is complete late 2025

The GRETA community eagerly awaits CD-1 from DOE. CD-1 allows real design (+ prototype, test, refine, ...) work to begin.

GRETINA was funded by the US DOE - Office of Science. Operation of the array at NSCL is supported by NSF under PHY-1102511(NSCL) and DOE under grant DE-AC02-05CH11231(LBNL).





#### Visit us at <u>http://greta.lbl.gov/</u>

http://gretina.lbl.gov/

Special thanks to these people for letting me show (butcher?) their slides:

#### **GRETINA/GRETA**

Augusto Macchiavelli Heather Crawford Paul Fallon

<u>Results</u> Michael Jones Simone Bottoni Lew Riley