

NUSTAR Collaboration Report Zsolt Podolyák

RRB, 17 May 2024



FAR News from the NUSTAR Board of Representatives



outgoing members





Andreas HeinzAnu Kankainen(Chalmers, Sweden)(JYFL, Finland)

spokesperson



Zsolt Podolyák (Surrey, UK)

new elected members



Lola Cortina Michael Block (Valencia, Spain) (Mainz, Germany)

Deputy spokesperson: Michael Block

https://faircenter.eu/user/experiments/nustar/nustarcollaboration/board-of-representatives

FAR News from the NUSTAR Council



new chair elect



Wolfram Korten (CEA, Saclay, France)

chair



Berta Rubio (IFIC, Valencia, Spain)

old chair elect (resigned)



Zsolt Podolyák (Surrey, UK)

Practically, there was a swap (through elections) between Wolfram Korten and Zsolt Podolyák

https://faircenter.eu/user/experiments/nustar/nustarcollaboration/council

FAR News from the NUSTAR Management Team



Technical Coordinator Resource Coordinator Technical Coordinator (a.D.)



Helena Albers (since 11/23)



Alexander Herlert (renewed 11/23)



https://faircenter.eu/user/experiments/nustar/nustarcollaboration/management-team



Timeline 2021 – 2029?



2021-2022FAIR-0 experiments

- 2023 No experiments (but there were tests)
- 2024-2025FAIR-0 experiments
- 2026-2027FAIR-0 experiments (G-PAC in 2025)
- End 2027 -> Early Science, then First Science

FAR The 11 Greatest Unanswered Questions of Physics



- 1. What is dark matter?
- 2. What is dark energy?

3. How were the heavy elements from iron to uranium made?

- 4. Do neutrinos have mass?
- 5. Where do ultrahigh-energy particles come from6. Is a new theory of light and matter neededto explain what happens at very high energiesand temperatures?
- 8. Are protons unstable?
- 9. What is gravity?
- 10. Are there additional dimensions?
- 11. How did the universe begin?



https://www.discovermagazine.com/the-sciences/the-11-greatest-unanswered-questions-of-physics



Overarching physics case: the creation of the (heavy) chemical elements



Big physics question requiring information on:

Equation of State Limits of existence Lifetimes, Masses P_{xn} values Fission Reactions in star environments



FAR One science, different observables, instrumentations



> 1000 listed "interested" scientists

- > 630 registered members
 (incl. students, etc.)
- ~ 420 senior members
 (PhD holder w/o Russia)
- > 150 institutes from 36 countries





- Very successful experimental campaigns in 2021 and 2022
 - Almost all experiment could be performed, not always with expected beam intensities (Pb)
- Summary of experiments
 - DESPEC: 27 days
 - ILIMA: 17 days
 - R3B: 44 days
 - S-FRS EC: 29 days
 - SHE: 90 days
 - Total: 113 days (SIS18) + 94 days (UNILAC)
- Phase-0 is still very productive
 - First technical and scientific papers published, many more are in preparation!
 - Essential to test detectors and to keep NUSTAR active within the community!
- In **2023** only engineering beam time, and NUSTAR detector tests.
- Phase-0 beamtimes restarted in February 2024.



Highlights and plans of the NUSTAR experiments





FAIR-0: ¹²C+¹²C benchmark case



A benchmark case to provide very precise data that minimise the uncertainty asociated to the reaction models

Exp. precision +-0.387%

Determined by the called Transmission method

 $\sigma_R = \sigma_{inel} + \sigma_I$ $\sigma_R = -\frac{1}{N_t} ln\left(rac{R_i}{R_o}
ight)$

 $R_{i/o}$ is the ratio of non-interacting nuclei after target and incoming nuclei for target in (out)



L. Ponath et al., submitted to PLB



Fission





measured Z (red) and A (green) fission yields in actinides and pre- actinides together with the neutron-rich (dark blue circles and dots) and other nuclei (light blue circles) that can be investigated at FAIR

FAIR FAIR-0: Prolate-oblate shape transition at ¹⁹⁰W₁₁₆





Hybrid array for fast-timing measurements





Super-FRS EC

New beam: ¹⁷⁰Er

First used in April 2024 New beams => New opportunities







Quasi-real-time range monitoring in hadron therapy using positron emitters of carbon and oxygen



FAIR E

Best candidate?

- ¹⁶O beam -> ¹⁵O 43 mb 122 s ¹⁴O 1.2 mb 71s
- ¹²C beam -> ¹¹C 47 mb 1221 s ¹⁰C 4.3 mb 19 s

Sivaji Purushothaman et al., Sci Rep 13, 18788 (2023) Quasi-real-time range monitoring by in-beam PET: a case for ¹⁵O

Quasi-real-time range monitoring Comparison of therapy relevant positron emitters of oxygen and carbon Time structure from experiment (n=1) If beam time structure was same (n=1) 140 & 150 [A.U.] Decay yield [A.U.] 0.75 0.75 10C & 11C Beam OFF: 1.5 s yield 140 & 150 0.5 0.5 Beam ON: 1 s 10C & 11C Decay Beam OFF: 1.5 s Beam ON: 2 s 0.25 0.25 Beam OFF: 2.8 s 10C 14O 15O 11C 10C 14O 15O 11C n= 500 n= 50 Decay yield [A.U.] Decay yield [A.U.] 0.75 0.75 10C & 11C 140 & 150 0.5 0.5 Beam ON: 1 s Beam OFF: 1.5 s 0.25 0.25 10C 14O 15O 11C 10C 14O 15O 11C

BIOMAT

(also hadron therapy with ¹¹C on mouse in Feb. 2024)

Rate capability test of the Cryogenic Stopping Cell





- New result of 2-10⁵ ions/s surpasses previous rate capability limit of 10⁴ ions/s is by more than 1 order of magnitude
- It validates the simulation model and provides additional support for the selected design concepts

FAIR GmbH | GSI GmbH



FAIR-0: Gamov-Teller Strength at N=50 and the puzzle of ¹⁰⁰Sn mass





First direct mass measurements of ⁹⁸Cd and ⁹⁷Rh with the FRS Ion Catcher

¹⁰⁰Sn mass:

New results in discrepancy of ¹⁰⁰Sn Q_{EC} values (Hinke et al. [1] and Lubos et al. [2])

 In recent work Mougeot et al. [3] derive the mass of ¹⁰⁰Sn from mass measurements of ⁹⁹⁻¹⁰¹In and published ¹⁰⁰Sn Q_{EC} values

 \rightarrow value of Hinke et al is favored

• This work:

Evolution of shifted two-neutron shell gap at N=50:

 \rightarrow Value of Hinke et al. [1] is favored.

Evolution of Gamov-Teller Strength at N=50:

 \rightarrow value of Lubos et al. [2] is favored.

Overall situation unclear, further experiments required

[1] C.Hinke et al., Nature **486** (2012) 341 [2] D.Lubos at al., PRL **122** (2019) 222502 [3] M.Mougeot et al., Nature Phys. **17** (2021) 1099 A. Mollaebrahimi et al., Phys.Lett. B 839, 137833 (2023)



FAIR-0: Two-phonon decay

Combined Schottky + Isochronous Mass Spectrometry

excitation energies down to ~ 100 keV and half-lives as short as ~ 10 ms. This week (⁹⁸Mo and ⁹⁸Zr) in May 2024 ^{Phys. Rev,}

D. Freire-Fernandez, W. Korten et al Phys. Rev, Lett., accepted

ILIMA

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Technical key points:

storage time (in hrs)

Lifetime of Highly Charge lons



- LOREX project: new solar neutrino detector
- Cosmochronometer of the s-process





ATLANTIS: the LaSpec beamline from Mainz, installed at ANL

Successful beamtimes in 2022, 2023: Ru + Pd

Using ²⁵²Cf fission beams

Planned experiments in 2024+: Tc, Ce, Nd, La, Rh

Using beams from neutron-induced fission on ²³⁵U targets

Beams of <100/s due to novel buncher & charge exchange cell







 $\Delta t < 1 \mu s$

eff.: ~80%





Cooler & Buncher RFQ

B. Maaß, K. König, L. Renth, P. Müller, W. Nörtershäuser, G. Savard et al.

Publication on Ru charge radii in preparation

SHE: Extending Laser spectroscopy to Fm (Z=100)



JGU



FAIR/GSI strategic operation scenario: ES, FS, towards FS+



Colours explanation (**Physics beamtime modes**)



FAIR Experiments location at Super-FRS (ES and FS)









-R3B starts moving into High-Energy Cave in 2026 => no R3B experiment in 2027 at FRS -Super-FRS commissioning upto FHF2 in 2027 (simple setup)

-Common NuSTAR experiment at FHF2

Experiments at FHF1 and FHF2 (also ESR via FRS, SHE etc.) this needs rearranging FHF1 while one area is used, the other cannot be accessed (radiation safety) => we need: beam in blocks (setting up between blocks and when beam not to Super-FRS)

-When R3B target area installed, only R3B experiments at FHF2 -DESPEC and Super-FRS EC experiments at FHF1



Start of FAIR: First NUSTAR experiment



Main aim to show that FAIR is running \Rightarrow Need to be **published fast** \Rightarrow Low risk (follows directly from SuperFRS commissioning) \Rightarrow Use some new capability: secondary beam intensity from primary beam from transmission higher beam energy (> 1 GeV/u) higher SuperFRS transmission Exps.: May 2022 first exp First publications: equipment

Lessons from FRIB (and RIKEN):

PRL on new lifetimes N>28 (published Nov. 2022) PRL on unexpected isomer 32Na (June 2023) PRL on new isotopes 198Pt beam (Feb. 2024; exp Feb. 2023)





FAIR Phase-0 is productive and assures readiness for Early and First Science

NUSTAR ES&FS experiments will mark the start of FAIR Exciting science opportunities based on SuperFRS and SIS100 (aligned with NuPECC LRP) Overarching physics case: creation of the chemical elements

Optimisation of the FAIR injector chain already for Early Science in parallel to FAIR construction (beam intensities!)

Vision for the completion of Super FRS low-energy branch (FS++) and ring branch (MSV)? GSI POF-5 strategy is being defined until 2034!



Thank you

FAIR RRB, 17 May 2024

