



Femtoscopy studies with HADES

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Femtoscopy NS, NSM FSI studies: NY, K_0^S , clusters, ...

Physics opportunity with proton beams at SIS100, Wuppertal, February 6-9, 2024



Femtoscopy

... the method to probe **geometric** and **dynamic** properties of the source (emission region, range of correlations-interactions, phase-space cloud, ...)

Classic femtoscopy

2**R**

Femtoscopy (originating from HBT):

the method to probe **geometric** and **dynamic** properties of the source

Space-time properties $(10^{-15}m, 10^{-23}s)$ determined thanks to two-particle correlations: **Quantum Statistics** (Fermi-Dirac, Bose-Einstein); **Final State Interactions** (Coulomb, strong)

determined assumed measured

$$C(k^*, r^*) = \int S(r^*) |\Psi(k^*, r^*)|^2 d^3r^* = \frac{Sgnl(k^*)}{Bckg(k^*)}$$

 k^* - momentum of the first particle in PRF

 r^* - Separation between emission points

 $S(r^*)$ – source function

 $\Psi(k^*, r^*)$ – two-particle wave function (includes e.g. FSI interactions)

 $\frac{Sgnl(k^*)}{Bckg(k^*)}$ – correlation function

Gateway to study interactions

If we assume we know the **source function**, measured **correlations** are used to determine **interactions in the final state**.

Space-time properties $(10^{-15}m, 10^{-23}s)$ determined thanks to two-particle correlations: **Quantum Statistics** (Fermi-Dirac, Bose-Einstein); **Final State Interactions** (Coulomb, strong)

assumed determined measured $C(k^*, r^*) = \int S(r^*) |\Psi(k^*, r^*)|^2 d^3 r^* = \frac{Sgnl(k^*)}{Bckg(k^*)}$

 $S(r^*)$ – source function

2**R**

 $\Psi(k^*, r^*)$ - two-particle wave function (includes e.g. FSI interactions) $\frac{Sgnl(k^*)}{Bckg(k^*)}$ - correlation function



Neutron stars Neutron star mergers



https://www.researchgate.net

Scanning various μ_B

.. to study strongly interacting matter .. to explore unknown QCD territory

Neutron star mergers CBM and HADES future



Temperature T<50 MeV

Density $n < 2 - 6n_0$

Reaction time $t \sim 10 \text{ ms}$ (GW170817)

Temperature T<10 MeV

Density $n < 10n_0$

Lifetime t ~ infinity

CBM and HADES future



Neutron star puzzle

- Hyperons: expected in the core of neutron stars; conversion of N into Y energetically favorable.
- Appearance of Y: The relieve of Fermi pressure → softer EoS → mass reduction (incompatible with observation).
- The solution requires a mechanism that could provide the **additional pressure** at high densities needed to make the EoS stiffer.
- A few possible mechanisms, one of them:
- Two-body YN & YY interactions
- A lot of experimental and theoretical effort to understand:
- The KN interaction, governed by the presence of $\Lambda(1405)$
- The nature of $\Lambda(1405)$, the consequences of KNN formation
- **K** and $ar{K}$ investigated to understand kaon condensation

 $M_{\rm NS} \approx 1 \div 2 M_{\odot}$ $R \approx 10-12 \text{ km}$ $\rho \approx 3 \div 5 \rho_0$



Neutron star puzzle



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Neutron stars (NS) are the remnants of the gravitational collapse of massive stars during supernova event.

Their masses and radii are of the order of $1-2~M_{\odot}$ and 10-12 km, respectively.

Central densities in the range of 4 – 8 times the normal nuclear matter saturation density, $\epsilon_0 \sim 2.7 \times 10^{14} \text{ g/cm}^3$ ($\rho_0 \sim 0.16 \text{ fm}^{-3}$)

Best suitable theory takes hyperons into account, Hyperons are expected to appear in the core of NS at ρ \sim 2- 3 ρ_0

Hyperons soften the EoS —> Reduction on maximum NS mass

Observation of the NS with $M_G > 2M_S$ is incompatible with such soft EoS

Although the existence of hyperons is energetically favorable, their existence makes the EoS softer and is not consistent with the experimental results. This is the essence of the **hyperon puzzle**.



Final State Interactions (+ other studies)

YN and YY interactions

• Experiment: More ... and more! interest about YN and YY interactions!



- Theory: Major steps forward have been taken (Lattice QCD).
- Numerous theoretical predictions exist, many experimental searches look for evidence for bound states.
- The existence of **hypernuclei** (confirmed by attractive YN interaction) → indicates the possibility to bind Y to N.
- The measurement of the YN and YY interactions leads to important implications for the possible formation of **YN** or **YY bound states**.
- A precise knowledge of these interactions help to explore unknown structure of neutron stars.

HADES spectrometer

- SIS-18 beams: protons (1-4 GeV), nuclei (1-2 AGeV), pions (0.4-2 GeV/c) secondary beam
- rare probes: (e⁺, e⁻), strangeness: K^{+/-,0}, Λ , Ξ ⁻, ϕ
- PID : $\pi/p/K dE/dx$ (MDC) and TOF
- electrons : RICH (hadron blind)
- neutral particles: ECAL

Geometry :

• full azimuthal, polar angles 18° - 85°







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Particle identification

ToF RPC



Ag

Ag





Signal reconstruction (p- Λ correlations)



Lednicky-Lyuboshitz model

The normalized pair separation distribution (source function) $S(r^*)$ is assumed to be Gaussian,

$$S(r^*)=(2\sqrt{\pi}r_0)^{-3}e^{-rac{r^{*2}}{4r_0^2}},$$
 Ref : Lednicky, Richard & J. Nucl. Phys. (Engl. Trac

Ref : Lednicky, Richard & Lyuboshits, V.L.. (1982). Sov. J. Nucl. Phys. (Engl. Transl.); (United States). 35:5.

The correlated function can be calculated analytically by averaging Ψ^{s} over the total spin S and the distribution of the relative distances $S(r^{*})$

$$egin{aligned} C(k^*) &= 1 + \sum_S
ho_s [rac{1}{2} |rac{f^S(k^*)}{r_0}|^2 ig(1 - rac{d_0^S}{2\sqrt{\pi}r_0} ig) + rac{2\mathbb{R}f^S(k^*)}{\sqrt{\pi}r_0} F_1(Qr_0) - rac{\Im f^S(k^*)}{r_0} F_2(Qr_0)] \ with \ F_1(z) &= \int_0^z dx e^{x^2 - z^2} / z \ and \ F_2(z) &= (1 - e^{-z^2}) / z \ \end{aligned}$$

 f_0 and d_0 - parameters of strong interaction.

Theoretical correlation function (k^*) depends on: R, f_0 and d_0 .

f₀ - the scattering length, determines low-energy scattering.

The elastic cross section, σ_e , (at low energies) determined by the scattering length, $\lim_{k\to 0} \sigma_e = 4\pi f_0^2$

 d_0 - the effective range, corresponds to the range of the potential (simplified scenario - the square well potential.

Results: centrality dependence



Centrality	Systematic Uncertainty		
0 - 10 %	15.30 %		
10 - 20 %	15.49 %		
20 - 30 %	19.00 %		





Results: k_T dependence



Lednicky-Lyuboshitz model

Model		$f_0^{S=0}$ (fm)	$f_0^{S=1}$ (fm)	$d_0^{S=0}$ (fm)	$d_0^{S=1}$ (fm)	n_{σ}
ND [77]		1.77	2.06	3.78	3.18	1.1
NF [78]		2.18	1.93	3.19	3.358	1.1
NSC89 [79]		2.73	1.48	2.87	3.04	0.9
NSC97 [80]	а	0.71	2.18	5.86	2.76	1.0
	b	0.9	2.13	4.92	2.84	1.0
	с	1.2	2.08	4.11	2.92	1.0
	d	1.71	1.95	3.46	3.08	1.0
	e	2.1	1.86	3.19	3.19	1.1
	f	2.51	1.75	3.03	3.32	1.0
ESC08 [81]		2.7	1.65	2.97	3.63	0.9
χEFT	LO [25]	1.91	1.23	1.4	2.13	1.8
	NLO [26]	2.91	1.54	2.78	2.72	1.5
Jülich	A [82]	1.56	1.59	1.43	3.16	1.0
	J04 [83]	2.56	1.66	2.75	2.93	1.4
	J04c [83]	2.66	1.57	2.67	3.08	1.1



parameter scan boundaries : f_0 [0.01, 5.0], d_{0s} [0.01, 2.0] and d_{0t} [0.01, 5.0]

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Lednicky-Lyuboshitz model Singlet Triplet 5 <mark>–</mark> \mathbf{q} HADES ² م HADES 10² 4.5 $|\Delta\chi^2|=1$ work in progress work in progress 1.8 4 1.6 3.5 1.4 3 1.2 2.5 1 2 10 0.8 1.5 0.6 0.4 **Extracted parameters : Extracted** parameters : 0.5 0.2 f_{0s} = 1.89 fm, d_{0s} = 3.76 fm $f_{0s} = 0.78$ fm, $d_{0s} = 0.01$ fm 00 00 0.6 0.8 1 1.2 1.4 1.6 1.8 1.5 2.5 3.5 4.5 0.2 2 3 0.4 0.5 4 2 1 5 f_{Ot} f_{0s} (بلا 1.8 0 1.7 $\lambda = 0.74$ $f_{0s} = 0.80^{+0.39}_{-0.32}$ HADES work in progress 1.6 $d_{0s}=0.01$ 1.5 $f_{0t} = 1.89^{+0.10}_{-0.09}$ 1.4 1.3 $d_{0t} = 3.76^{+0.27}_{-0.25}$ 1.2 **1.1**₿ $r = 2.24^{+0.12}_{-0.11}$ è te 0.9 ⊨ 0.05 0.15 0.2 0.25 0.1 0.3 0

k* (MeV/c)

Parameters scan



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Femtoscopy with neutral kaons, planned at HADES



 $q_{inv} = \sqrt{(\vec{p_1} - \vec{p_2})^2 - (E_1 - E_2)^2}$ 4 parameterizations of SI FSI, scanning of parameters welcome!

Some examples from ALICE



CATS: EPJA 78 (2018) Projector: EPJC 82 (2022) Review 1: Prog.Part.Nucl.Phys. 112 (2020) Review 2: Ann. Rev. Nucl. Part. Sci. 71 (2021) p-*ϕ* bound state: arXiv:2212.12690 p-K: PRL 124 (2020) 092301 p-K: PLB 822 (2021), EPJC (2022) p-p, p-Λ, Λ-Λ: PRC 99 (2019) 024001 A-A: PLB 797 (2019) 134822 p-Ξ-: PRL. 123 (2019) p-Ξ-, p-Ω : Nature 588 (2020) 232–238 p-Σ⁰: PLB 805 (2020) 135419 p-φ: PRL 127 (2021) $p - \bar{p}, \Lambda - \bar{\Lambda}, p - \bar{\Lambda}$: PLB 829 (2022) р-Л: PLB 832 (2022) 137272 Λ – Ξ: PLB 137223 (2022) D-p: PRD 106, 052010 (2022) ppp, ppA: arXiv:2206.03344

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Some examples from ALICE

nΣŪ

2

2.5

 m_{τ} (GeV/ c^2)

0.8

1.5



CATS: EPJA 78 (2018) Projector: EPJC 82 (2022) Review 1: Prog.Part.Nucl.Phys. 112 (2020) Review 2: Ann. Rev. Nucl. Part. Sci. 71 (2021) p-*φ* bound state: arXiv:2212.12690 p-K: PRL 124 (2020) 092301 p-K: PLB 822 (2021), EPJC (2022) p-p, p-A, A-A: PRC 99 (2019) 024001 Λ-Λ: PLB 797 (2019) 134822 p-Ξ-: PRL. 123 (2019) p-Ξ-, p-Ω-: Nature 588 (2020) 232–238 p-Σ⁰: PLB 805 (2020) 135419 p-q: PRL 127 (2021) $p - \bar{p}, \Lambda - \bar{\Lambda}, p - \bar{\Lambda}$: PLB 829 (2022) р-Л: PLB 832 (2022) 137272 Λ – Ξ: PLB 137223 (2022) D-p: PRD 106, 052010 (2022) ppp, ppA: arXiv:2206.03344

For lower collisions energy the fraction of pure correlation (contamination of feed-down contribution, residual correlations) is higher that at LHC.

Some examples from ALICE



E-M probes with ECAL



Light nuclei (clusters) production at HADES



0

20

40

k*[MeV/c]

140

p-d: T.C. Black et al., Phys.Lett.B 471, 103 (1999)

Summary

A. Sorensen, HZ et al. <u>arXiv:2301.13253</u> [nucl-th] Prog.Part.Nucl.Phys. 134 (2024) 104080



Femtoscopy is a perfect tool to study two- and more particle interactions

Thank you

Smoothness approximation (from Scott Pratt)



See S.Pratt talk during WPCF 2023

https://agenda.infn.it/event/33324/contributions/211990/attachments/112571/161133/wpcf_2023.pdf

Smoothness approximation (from Scott Pratt)

Both use same source function:



Approximations

The Truth:

$$P(p_a,p_b) = \sum_{f'} \left| \int dx_a dx_b T_{f'}(x_a,x_b) \phi_{f'}(x_a,x_b;p_a,p_b,)
ight|^2$$

Sum over all "remainder" states f'

APPROXIMATIONS

1) $\phi(x_a, x_b; p_a, p_b)$ does not depend on f'

 a) fails if phase space density is high (identical particles) multi-particle symmetrization is important otherwise, must calculate for all momenta, then integrate over all other particles

 b) fails if interaction with other particles lasts long time at sufficiently small relative momentum, this is fine Coulomb with other particles slowest other interaction

$$\phi_f(x_a, x_b; p_a, p_b)
ightarrow \phi(x_a - x_b; p_a, p_b)$$

After first approximation:

$$P(p_a,p_b)=\sum_{f'}\left|\int dx_a dx_b T_{f'}(x_a,x_b)\phi(x_a,x_b;p_a,p_b,)
ight|^2$$

APPROXIMATIONS

2) Emission (T-matrix) is independent.

Sum over f' and T-matrices must factorize

Ignores other correlations (energy/momentum/charge) conservation...

good at small relative momentum (other sources have longer characteristic scales)

$$\sum_{f'}
ightarrow \sum_{f'_a} \sum_{f'_b}, \ T_{f'_a}(x_a, x_b)
ightarrow T_{f'_a}(x_a) T_{f'_b}(x_b)$$

Approximations

After approximations 1 & 2:

Define:

$$s_a(x,p)\equiv\sum_{f_a'}\int d\delta x\;T^*_{f_a'}(x+\delta x/2)T_{f_a'}(x-\delta x/2)e^{ip\cdot\delta x}$$
 is given

this gives

$$P_{ab}(p_a,p_b) = \int dx_a dx_b d\delta x d ilde{q} s_a (ar{P}_a + ar{q},x_a) s_b (ar{P}_b - ar{q},x_b)
onumber \ e^{i ilde{q}\delta x} \phi_q^*(x_a - x_b + \delta x/2) \phi_q^*(x_a - x_b - \delta x/2)$$

3) Smoothness approximation:

a) Ignore \tilde{q} dependence in $s_a(\bar{P}_a + \tilde{q}, x_a)$ and $s_b(\bar{P}_b - \tilde{q}, x_b)$, b) replace $s_a(\bar{P}_a + \tilde{q}, x_a)s_b(\bar{P}_b + \tilde{q}, x_b)$ with $s_a(p_a, x_b)s_b(p_b, x_b)$ or $s_a(E_a, \vec{p}_{a,cm}s_b(E_b, \vec{p}_{b,cm})$

- good when emissions are thermal and matrices are broad
- questionable if relative momentum is small
- necessary if you don't know off-shell behavior of s(p, x)
- for coalescence you can add $e^{B/T}$ factor

Last approximation: non-simultaneous wave functions

$$\phi_q(x_1-x_2)=\phi(\Delta t=0,ec x_1-ec x_2)$$
 in pair frame

4) Non-simultaneous emission

- no problem for pure HBT

- should be fine for small relative momentum

Sometimes interactions involve change of degrees of freedom: No interaction through potential

Examples: $K^+K^- \leftrightarrow \phi$ $\alpha d \leftrightarrow^6 \text{Li}$

Wavefunction paradigm questionable — but thermal equilibrium still applies

When are the approximations good?

Femtoscopy:

- Emission uncorrelated aside from FSI
- Relative motion is small, $q/\mu \lesssim 0.1$
- Phase space density not high (as long as phase space densities ≤ 0.5)
- Range of interaction smaller than source size
- Rearrangement interactions, e.g. $K^+K^- \leftrightarrow \phi$ where wave function paradigm is questionable

Coalescence:

- Same as above
- Wave function should not have high p components (low B)
- Should correct for binding energy: $e^{B/T}$

<u>Thermal</u>:

- Must be at freeze out!!
- Whenever wave function extent is < < source size
- OK with rearrangement interactions

Central H.I. Collisions: – Usually very solid *pp* Collisions: – Be more careful Rearrangment interactions – Be careful

Signal reconstruction



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- Distance of closest approach (DCA) between the daughter tracks and the primary vertex,
- \rightarrow Dau1VD = > 8 mm
- \rightarrow Dau2VD = > 24 mm
- DCA between reconstructed mother
 track and primary vertex (Mot-VD) = < 5
 mm
- Distance between the primary and secondary vertex (VDX) = > 65 mm
- DCA between the two daughter tracks
 (MTD)
 = < 6 mm
- Opening angle between the two daughter tracks (A) = > 15 $^{\circ}$

Lednicky-Lyuboshitz model



Lednicky & Lyuboshitz analytical model

Calculate correlation function

fit

Experimental correlation function

x² calculation minimum determination

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1. The Lednicky-Luboshitz semi-analytical model (utilized in CorrfitCumac codes) provides an immediate correlation function value but may be computationally intensive due to integral calculations.

2. The first fitter employs ROOT minimizers, offering precise statistical uncertainty estimation, but it operates on "continuous" maps with limited control over parameter steps.

3. The second fitter, Hal:Minimizer, accommodates "non-continuous" functions, allowing parameters to change in discrete steps. However, it provides only approximate uncertainty estimates.



YN interactions at STAR

